PPC 2016

N-Relaxion: Large field excursions from a few site relaxion model

Ricardo D'Elia Matheus



The Hierarchy Problem

Think about the Standard Model (SM) as an EFT with a cut-off at M_p:

$$V(H)=m_H^2(lpha,eta)H^2+\lambda h^4+\mathcal{O}(1/M_p^2)$$

$$\langle H
angle = v$$

The only mass scale is M_p!

Technical Naturalness All dimensionless Wilson coefficients should be of order one.

$$m_H^2 \equiv
ho \; M_P^2$$

$$Dim[
ho]=Dim[\lambda]=0$$

$$\rho \approx \lambda \approx 1$$

Important exception: taking a coefficient to zero increases symmetry. In that case it can be arbitrarily small.

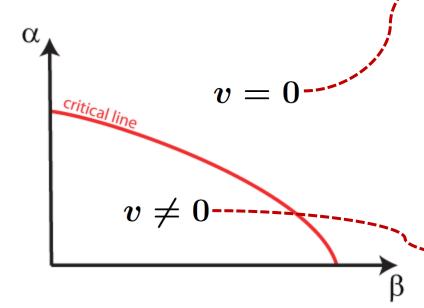
The Hierarchy Problem

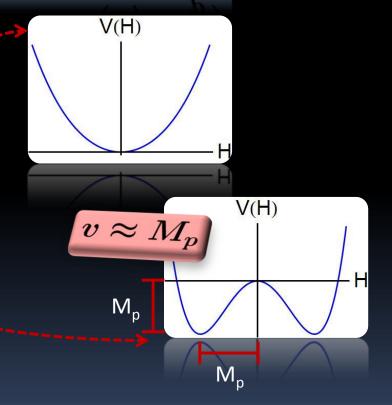
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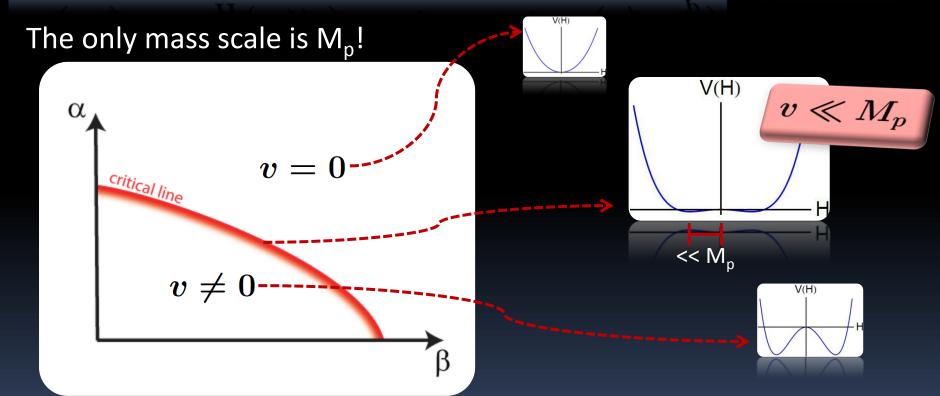


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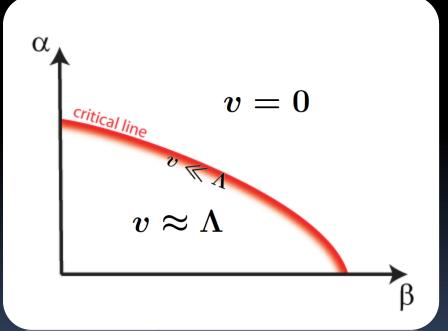
inspired by Alex Pomarol)

Solving the Hierarchy Problem

Question: how come we live so close to the line?

Two answers: (1) Some symmetry forces it! (SUSY)

(2) The cut-off Λ , is not really M_p . In fact $\Lambda << M_p$ and $\Lambda \sim 1$ TeV (Composite Models, Extra Dimensions et al.)



Both **DEMAND** new physics @ ~TeV

A New Hope

Question: how come we live so close to the line?

The Third Way: (3) History! Make α and β dynamical (fields in fact)

(stupid) Example: $m_H^2(lpha,eta)H^2 o lphaeta H^2$

$$m_H^2 = \langle lpha
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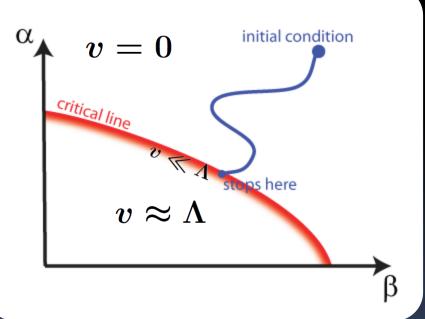
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But how does the evolution stop?

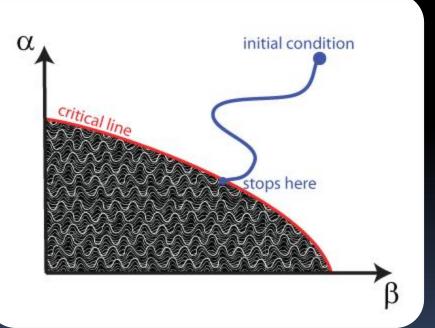
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But how does the evolution stop?

Local Minima! A whole LOT OF local minima!

Can it be done in a (technically) natural way?

(spoiler: yes! But...)

Introduce one scalar field ϕ , and:

 $v \approx \Lambda$

 $\phi_c \equiv \Lambda/g$

$$m_H^2 o m_H^2(\phi) = -\Lambda^2 \left(1 - \frac{g\phi}{\Lambda}\right)$$

"rolls" down

Must stop here...

Large Field Excursions!

 $|g\ll 1|$

 $\phi pprox \Lambda/g \gg \Lambda$

The minimal model:

$$V(\phi,H)=\Lambda^3 g\phi-rac{1}{2}\Lambda^2igg(1-rac{g\phi}{\Lambda}igg)H^2+\epsilon\Lambda_c^2H^2\cos(\phi/f)$$
 Linear slope for ϕ 1/2 $m_{
m H}^2$ Local minima in ϕ

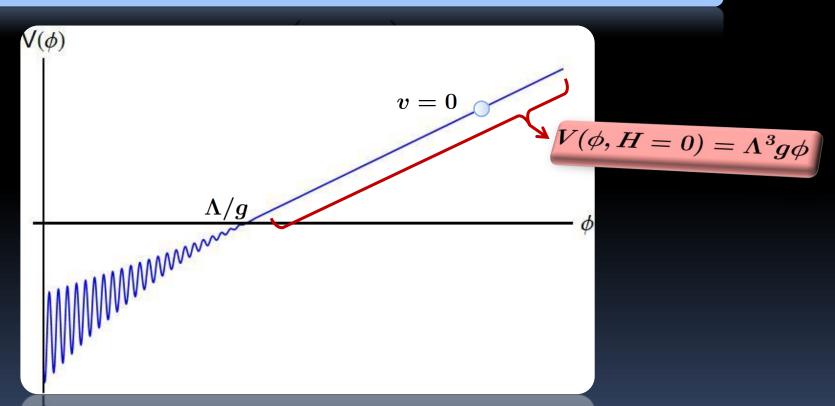
Both g and ε break shift symmetries (more about that later) and can be naturally small!

 Λ is the cut-off for the SM

 $\Lambda_{\rm c}$ is the scale at which the periodic potential is generated

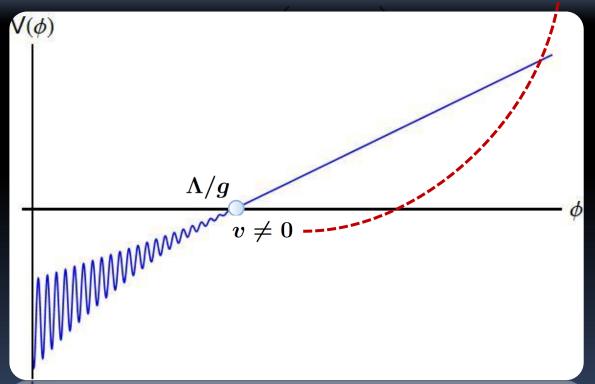
The minimal model:

$$V(\phi,H) = \Lambda^3 g \phi - \frac{1}{2} \Lambda^2 \bigg(1 - \frac{g \phi}{\Lambda} \bigg) \, H^2 + \epsilon \Lambda_c^2 H^2 \cos(\phi/f)$$



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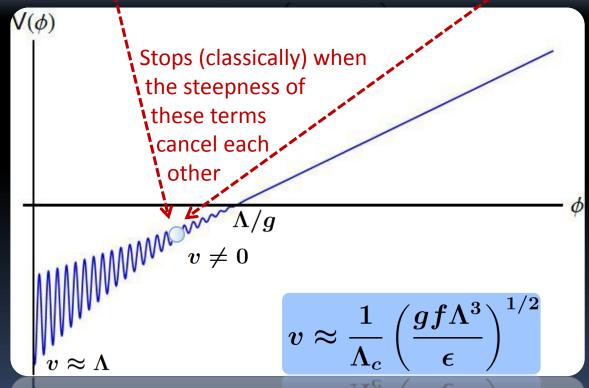
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Becomes more important as *v* grows

The minimal model:

$$V(\phi,H) = \Lambda^3 g \phi - \frac{1}{2} \Lambda^2 \bigg(1 - \frac{g \phi}{\Lambda} \bigg) \, H^2 + \epsilon \Lambda_c^2 H^2 \cos(\phi/f) \bigg)$$



The overall slope is controlled by g.

$$v \ll \Lambda$$

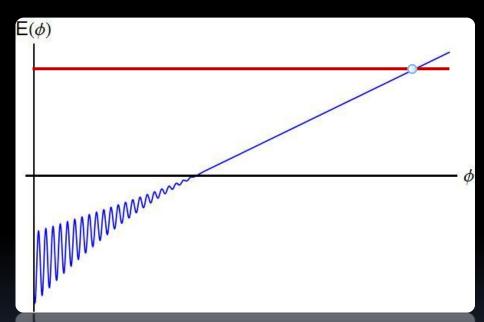
$$g\ll 1$$

Technically Natural!

NO NEW PHYSICS close to *v*

So far, so good. Now on to the little dirty details:

Do we risk overshooting? Do we need to start close to ϕ_c ?



NO, if slow rolling (during an inflationary epoch). Inflation introduces Hubble friction:

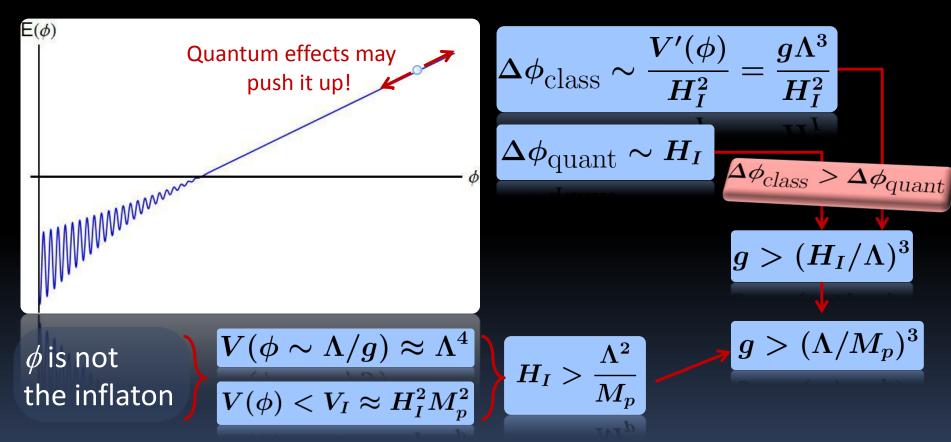
(\phi is not the inflaton)

Consequence: homework for cosmologists as a long period of inflation is needed ($N_e \sim 10^{40}$, smaller if H_l is not constant – Pattl, Schwaller, arXiv:1507.08649)

Optional approach: no inflation, temperature dependent potential. E. Hardy, arXiv:1507.07525

So far, so good. Now on to the little dirty details:

Limitations: Inflation → de Sitter space → Temperature (from Horizon)



So far, so good No

Limi†

 $E(\phi)$

 $g < 1 \rightarrow \Lambda$ is not M_p

g cannot be arbitrarily small

(otherwise it will bring Λ to the TeV scale and

below)

 ϕ is not the inflaton

$$\left[V(\phi\sim\Lambda/g)pprox\Lambda^4
ight]$$

$$V(\phi) < V_I pprox H_I^2 M_p^2$$

$$H_I > rac{\Lambda^2}{M_p}$$

(from Horizon)

$$rac{g\Lambda^3}{H_I^2}$$

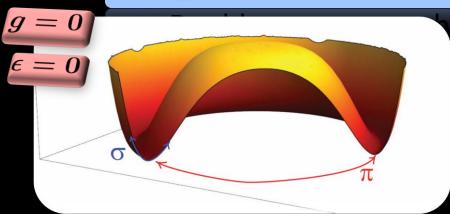
$$\Delta\phi_{class} > \Delta\phi_{quant}$$

$$g>(H_I/\Lambda)^3$$

$$g > (\Lambda/M_p)^3$$

Symmetries

$$\mathcal{L} = rac{1}{2} \partial_{\mu} \phi \partial^{\mu} \phi - \Lambda^3 g \phi - rac{1}{2} \Lambda g \phi H^2 - \epsilon \Lambda_c^2 H^2 \cos(\phi/f)$$

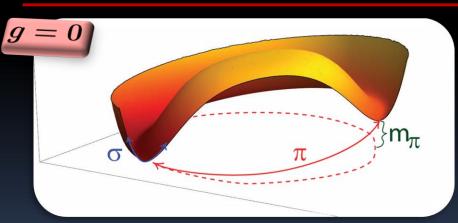


Nambu-Goldstone Bosons (NGB)

$$\mathcal{L}(\pi) = \mathcal{L}(\pi+c), orall c$$
 $V(\pi) = 0$

Compact Field Space $(2\pi f)$

Continuous shift symmetry



Pseudo-NGB (pNGP)

$$\mathcal{L}(\pi) = \mathcal{L}(\pi + 2n\pi f)$$

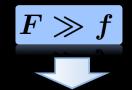
Compact Field Space $(2\pi f)$

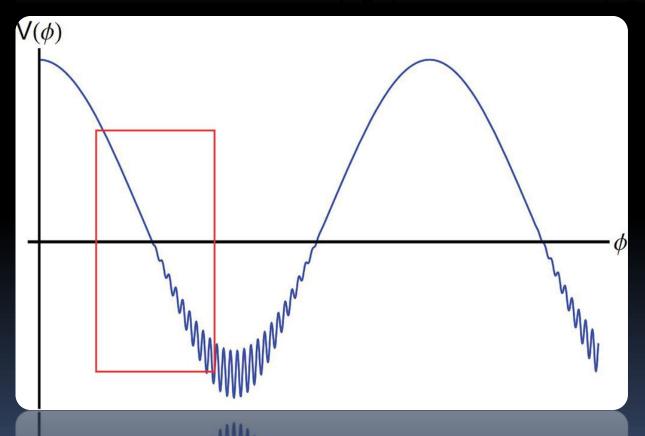
Discrete shift symmetry

Allowed potential MUST be periodic in the field!

Symmetries

$$V(\pi,H) \sim \kappa_1(H^2) \cos\left(rac{\pi}{F}
ight) + \kappa_2(H^2) \cos\left(rac{\pi}{f}
ight)$$





But how can we get the same pNGB to have two very different periods

(compact field spaces)?

Clockwork Relaxion

Key element: many pNGBs with the same decay constant f:

$$\mathcal{L}_{pNGB} = f^2 \sum_{j=0}^N \partial_\mu U_j^\dagger \partial^\mu U_j + \left(\epsilon f^4 \sum_{j=0}^{N-1} U_j^\dagger U_{j+1}^3 + h.c.
ight) + \cdots \ U_j \equiv e^{i\pi_j/(\sqrt{2}f)} \ U(1)^{\mathsf{N}+1} \qquad \qquad \mathsf{U}(1)^{\mathsf{N}+1}
ightarrow \mathsf{U}(1)^{\mathsf{N}+1}
ightarrow \mathsf{U}(1) = Q_j/3$$

$$V^{(2)} = rac{1}{2} \epsilon f^2 \sum_{j=0}^N (q \pi_{j+1} - \pi_j)^2$$

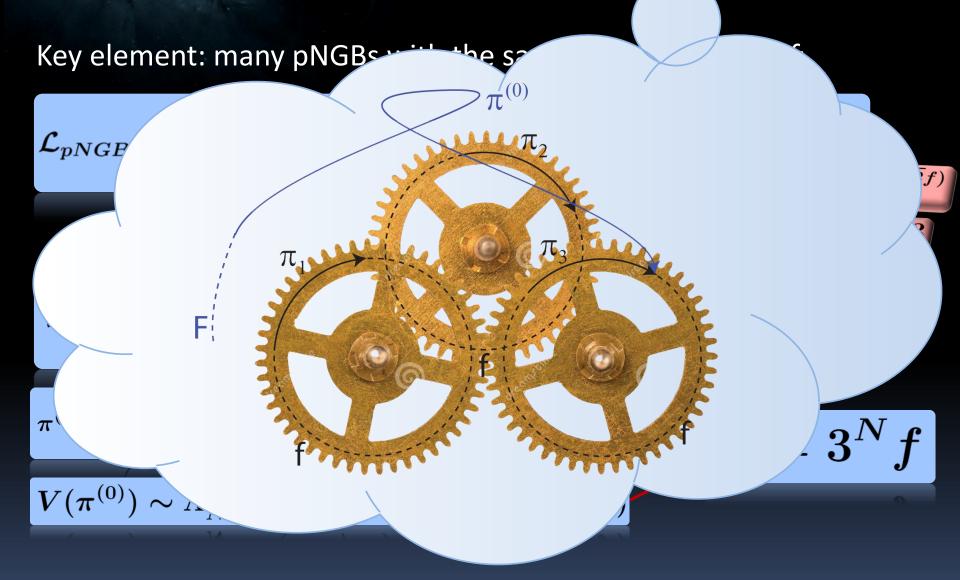
$$\pi^{(0)} \sim \left(\pi_0 + rac{1}{3}\pi_1 + rac{1}{9}\pi_2 + ... + rac{1}{3^N}\pi_N
ight)$$

$$V(\pi^{(0)}) \sim \Lambda_N^4 \cos(\pi^{(0)}/F) + \Lambda_0^4 \cos(\pi^{(0)}/f)$$

$$|F=3^N f|$$

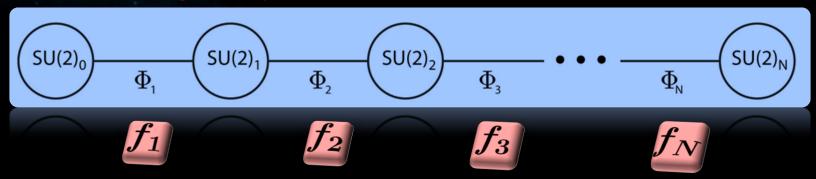
A Clockwork Axion Kaplan, Rattazzi,

Clockwork Relaxid



The N-site model & extra dimensions

Arkani-Hamed, Cohen, Georgi, arXiv:hep-th/0104005v1



$$S_4 = \int \, d^4x \, \left\{ -rac{1}{2} \sum_{j=0}^N {
m Tr}[F_{\mu
u,j} F_j^{\mu
u}] + \sum_{j=1}^N {
m Tr}[(D_\mu \Phi_j)^\dagger (D^\mu \Phi_j)] - V(\Phi)
ight\}$$

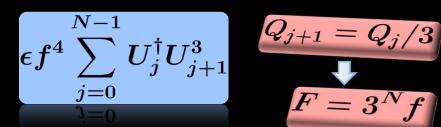
Large N limit: SU(2) gauge theory in five dimensions.

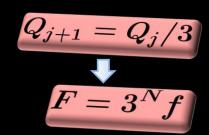
Choice of scales determines metric:

$$f_j = f, \;\; orall j \;\;
ightharpoonup \;\;$$
 Flat extra dimension

$$f_j = fq^j, \;\; 0 < q < 1$$
 $ightharpoonup$ AdS $_5$

Kaplan-Rattazzi clockwork axion:

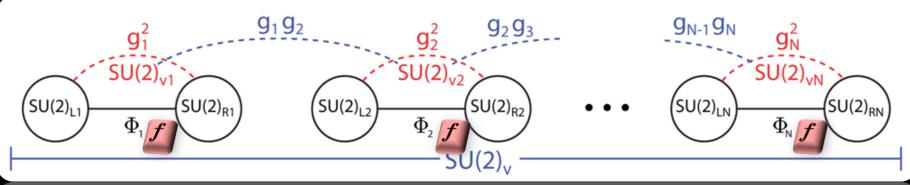




No continuum limit!

Goals:

- Find a model closer to a dimensional deconstruction that: (i) has a relaxion and (ii) provides a effective scale F much greater than f.
- Emulate the discretization of AdS₅ (which is motivated by dualities to strongly coupled theories). This is tricky, since we are taking $f_i = f$
- Generalize to non-abelian symmetries



$$\sum_{j=1}^{N} ext{Tr} iggl[\partial_{\mu} \Phi_{j}^{\dagger} \, \partial^{\mu} \Phi_{j} + rac{f^{3}}{2} (2 - \delta_{j,1} - \delta_{j,N}) g_{j}^{2} iggl(\Phi_{j} + \Phi_{j}^{\dagger} iggr) iggr] - rac{f^{2}}{2} \sum_{j=1}^{N-1} g_{j} g_{j+1} ext{Tr} \left[(\Phi_{j} - \Phi_{j}^{\dagger}) (\Phi_{j+1} - \Phi_{j+1}^{\dagger})
ight]$$



$$g_j \to q^j, \quad 0 < q < 1$$

$$q=rac{g_{j+1}}{g_j}$$

Small symmetry breaking parameters

(will be hierarchical to emulate AdS₅)



$$ec{\eta_0} = \sum_{j=1}^N rac{q^{N-j}}{\sqrt{\sum_{k=1}^N q^{2(k-1)}}} ec{\pi}_j$$

Same as the Wilson Line in AdS₅!

$$\mathcal{L}_{\eta} = \sum_{j=1}^{N} \left[rac{1}{2} \partial_{\mu} ec{\eta}_{0} \cdot \partial^{\mu} ec{\eta}_{0} + f^{4} (2 - \delta_{j,1} - \delta_{j,N}) q^{2j} \cos rac{\eta_{0}}{f_{j}}
ight] + \sum_{j=1}^{N-1} f^{4} q^{2j+1} \sin rac{\eta_{0}}{f_{j}} \sin rac{\eta_{0}}{f_{j+1}}$$

$$M_{\pi}^{2} = f^{2} \begin{pmatrix} q^{2} & -q^{3} & 0 & \dots & 0 & 0 \\ -q^{3} & 2q^{4} & -q^{5} & \dots & 0 & 0 \\ 0 & -q^{5} & 2q^{6} & \dots & 0 & 0 \\ \vdots & \vdots & \vdots & \dots & \vdots & \vdots \\ 0 & 0 & 0 & \dots & 2q^{2(N-1)} & -q^{2N-1} \\ 0 & 0 & 0 & \dots & -q^{2N-1} & q^{2N} \end{pmatrix}$$
(massless at tree level, loops induce: $m = f^{2}q^{2N}$)

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ight] + \sum_{j=1}^{N-1} f^{4} q^{2j+1} \sin rac{\eta_{0}}{f_{j}} \sin rac{\eta_{0}}{f_{j+1}}$$

New scales for oscilation

$$f_j \equiv f q^{j-N} {\cal C}_N$$
 ${\cal C}_N pprox { extbf{1}}$ (small q or big N)



$$igotharpoondown f_Npprox f$$

$$lacksquare F = f_1 pprox f/q^{N-1}$$

$$M_{\pi}^{2} = f^{2} \begin{pmatrix} q^{2} & -q^{3} & 0 & \dots & 0 & 0 \\ -q^{3} & 2q^{4} & -q^{5} & \dots & 0 & 0 \\ 0 & -q^{5} & 2q^{6} & \dots & 0 & 0 \\ \vdots & \vdots & \vdots & \dots & \vdots & \vdots \\ 0 & 0 & 0 & \dots & 2q^{2(N-1)} & -q^{2N-1} \\ 0 & 0 & 0 & \dots & -q^{2N-1} & q^{2N} \end{pmatrix}$$

$$(\text{massless at tree level, loops induce: } m = f^{2}q^{2N})$$

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ight] + \sum_{j=1}^{N-1} f^{4} q^{2j+1} \sin rac{\eta_{0}}{f_{j}} \sin rac{\eta_{0}}{f_{j+1}}$$

$$f_j \equiv f q^{j-N} \mathcal{C}_N$$
 $\mathcal{C}_N pprox 1$

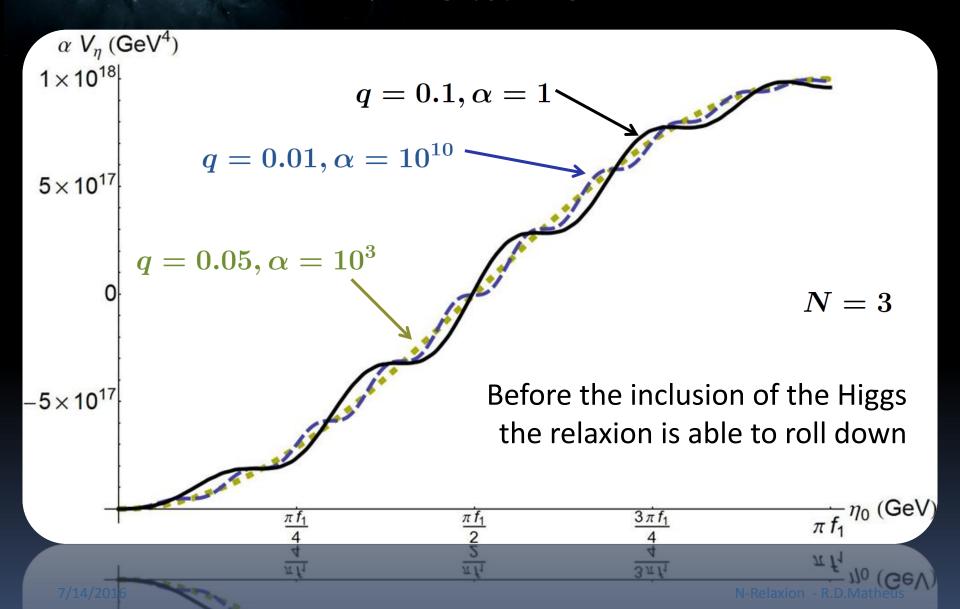


Amplitudes are also controlled by q Bigger frequencies ← smaller amplitudes (only the first few really matter)

$$lue{f_N}pprox f$$

$$lacksquare F = f_1 pprox f/q^{N-1}$$

 $V(\eta_0)$ gets flat for q << 1



Interaction with the Higgs:

$$\mathcal{L}_{\eta,H} = \left(1 + rac{|H|^2}{\Lambda^2}
ight)\mathcal{L}_\eta + |D_\mu H|^2 + rac{\Lambda^2}{2}|H|^2 - rac{\lambda_H}{4}|H|^4 + \epsilonrac{\Lambda_c}{16\pi}\mathrm{Tr}[\Phi_N + \Phi_N^\dagger]|H|^2$$

Most general thing you can do



New explicit breaking at site N

$$\left[\epsilon f^2 |H|^2 \cosrac{\eta_0}{f_N}
ight]$$

Generates the linear terms

$$-\Lambda^3 g\phi -rac{1}{2}\Lambda g\phi H^2$$

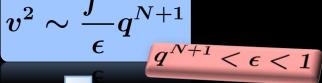
Generates high frequency oscillations once $v \neq 0$

- Also generates high frequency oscillations everywhere, double scanner needed!
- Modification of AdS₅ near the infrared brane (IR), enforces that the SM Higgs should be IR localized

Interaction with the Higgs:

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Solving for the classical stopping of the rolling: $v^2 \sim \frac{f^2}{q^{N+1}}$



Constraints:

"not the inflaton"

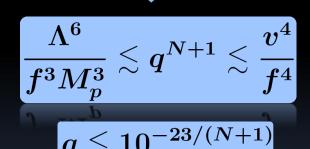
"classical rolling vs quantum

fluctuations"

$$q^{N+1}>rac{\Lambda^6}{f^3M_p^3}$$

"suppressing terms like $\varepsilon \cos^2$ "

$$\epsilon < v^2/f^2$$



$$f\lesssim 10^8~{
m GeV}$$

 $q^{N+1} < \epsilon < 1$

N-Relaxion

Interaction with the Higgs:

$$\mathcal{L}_{\eta,H} = \left(1 + rac{|H|^2}{\Lambda^2}
ight)\mathcal{L}_{\eta} + |D_{\mu}H|^2 + rac{\Lambda^2}{2}|H|^2 - rac{\lambda_H}{4}|H|^4 + \epsilonrac{\Lambda_c}{16\pi}\mathrm{Tr}[\Phi_N + \Phi_N^{\dagger}]|H|^2$$

Solving for the classical stopping of the rolling: $v^2 \sim \frac{f^2}{2} q^{N+1}$

Co
$$q=10^{-24/(N+1)}$$
 & $\epsilon=10^{-12}$ $fpprox 10^8~{
m GeV}$

$$N=2
ightarrow m_{\eta_0}pprox 10^{-7} {
m eV}$$
 $N=3
ightarrow m_{\eta_0}pprox 10^{-11} {
m eV}$

$$rac{\Lambda^6}{f^3 M_p^3} \lesssim q^{N+1} \lesssim rac{v^4}{f^4}$$

$$q\lesssim 10^{-23/(N+1)}$$

$$f\lesssim 10^8~{
m GeV}$$

Conclusions

- The relaxation models are a proof of concept. If we come to the conclusion that they are self-consistent, then the hierarchy problem ceases to be an argument for new physics at the TeV scale.
- We manage to build an N-site relaxion model with a well defined continuum limit. Some improvements are needed and/or interesting:
 - To build the double scanner sector (or another solution to the high frequency oscillations induced by the Higgs)
 - To explore other symmetry breaking patterns. Can any of the possible patterns allow us to increase the cut-off? Or do away with the double scanner?
 - What about the continuum limit? What theory do we get in AdS₅?

Thank You!