

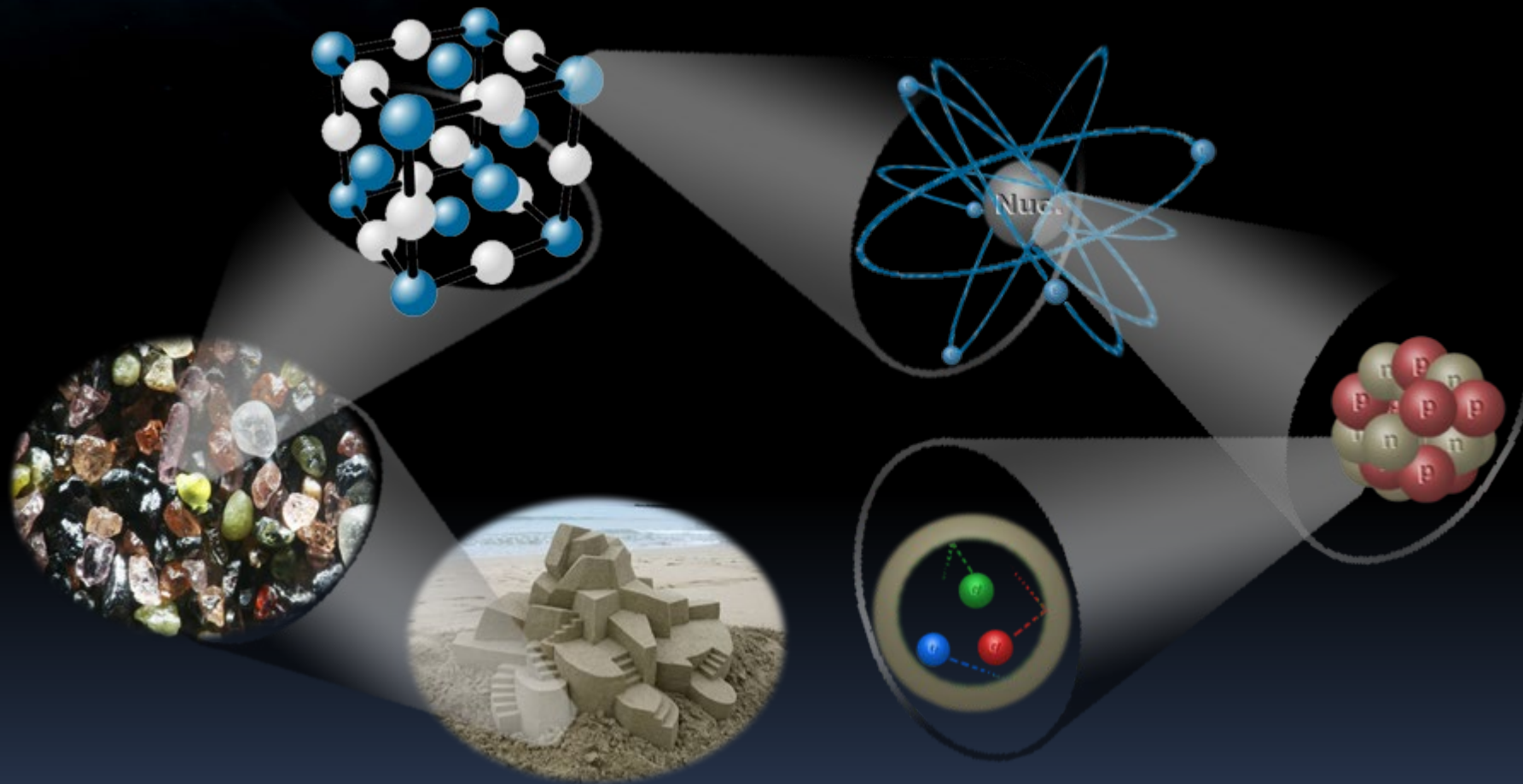
A dark, cosmic background image featuring a nebula or galaxy structure in the upper left corner, with a gradient from black to dark blue towards the bottom right.

INFIERI 2023

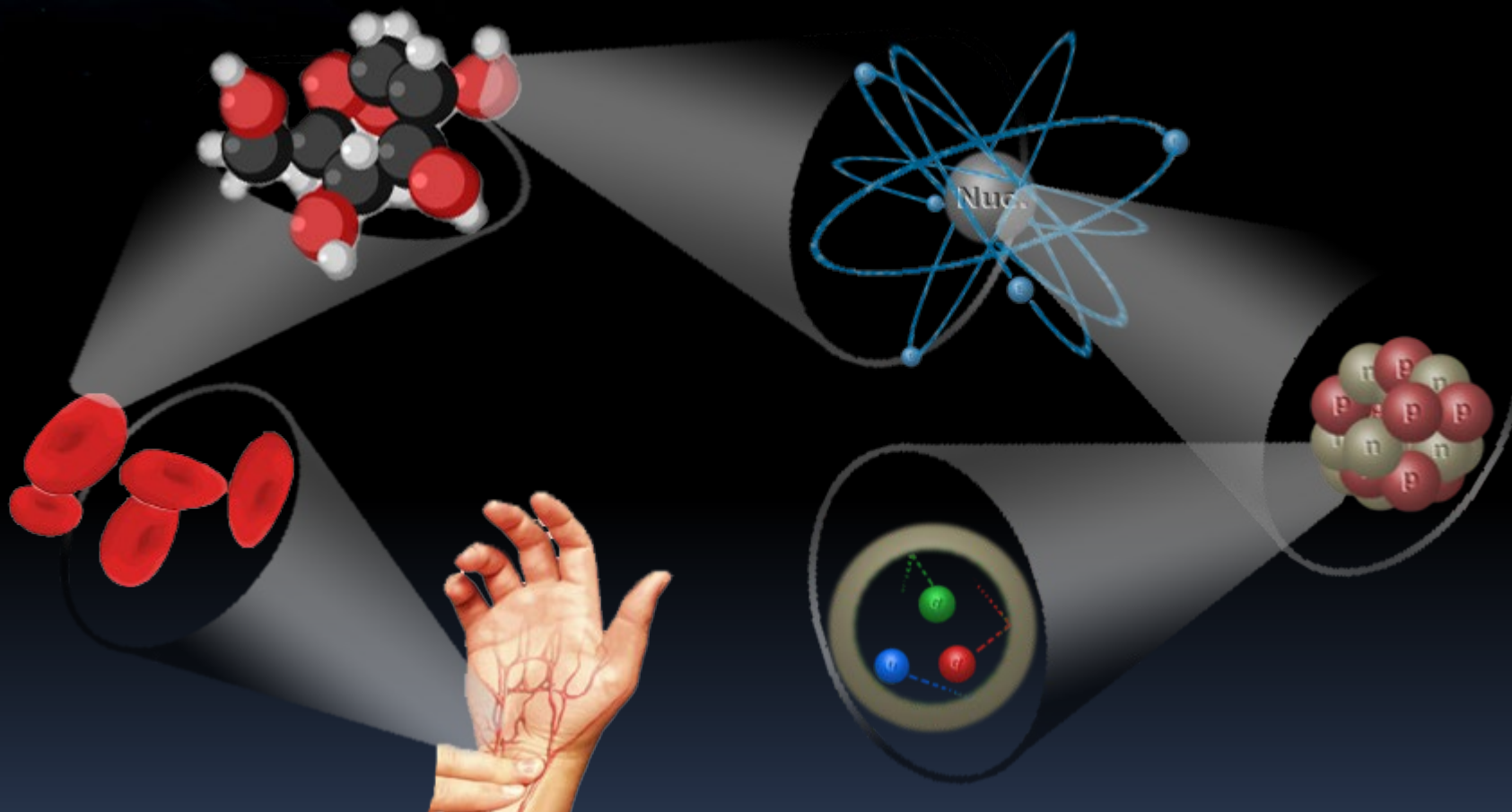
Masterclass: Particle Physics

Ricardo D'Elia Matheus

What do we want from Particle Physics?

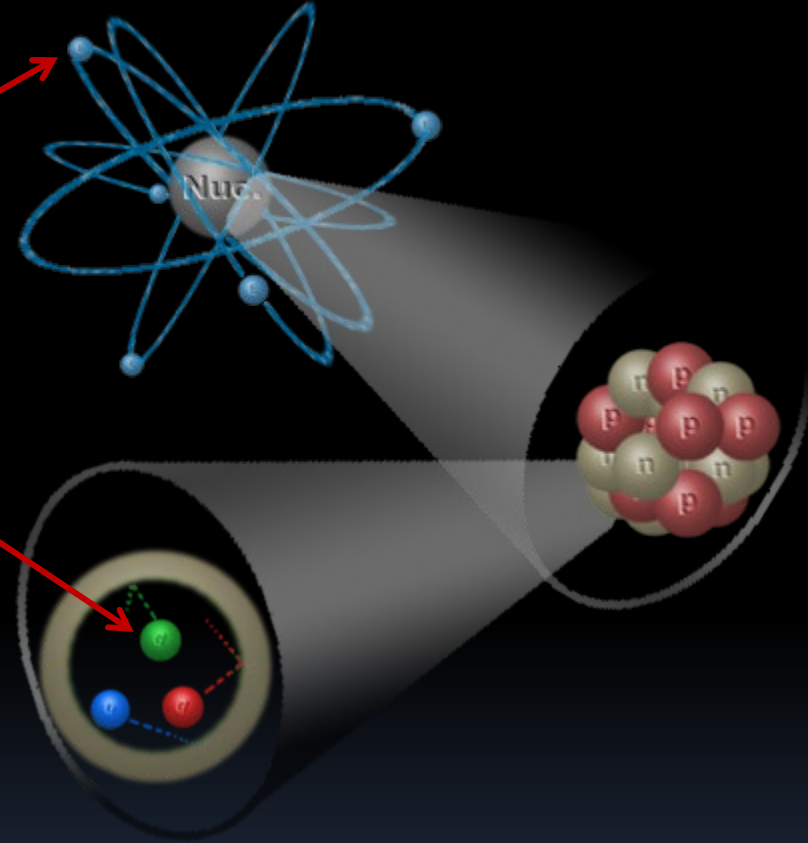


What do we want from Particle Physics?



What do we want from Particle Physics?

Well, but do “particles”
can really be
treated as “billiard balls”?

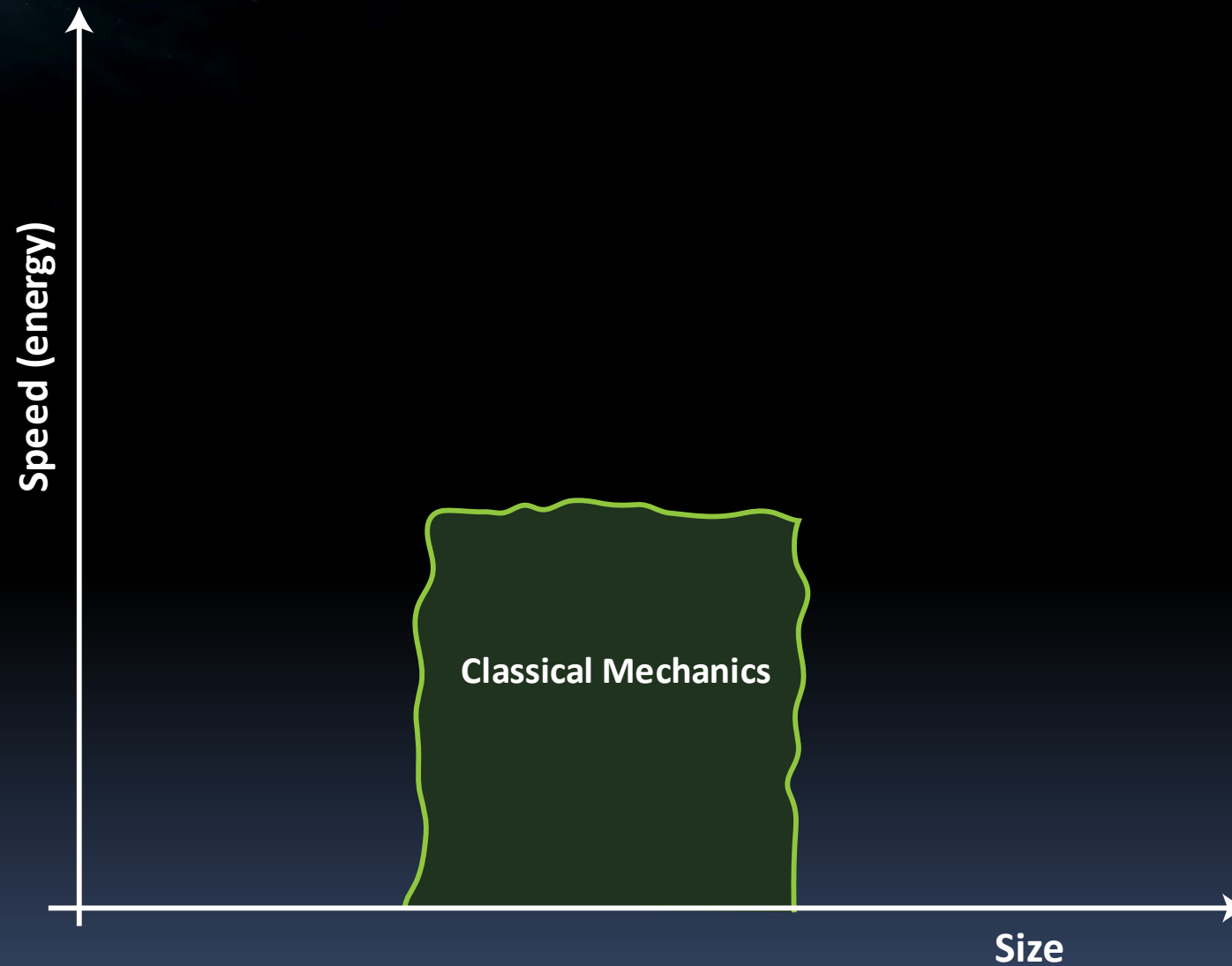


What is the appropriate mathematical description for these objects?

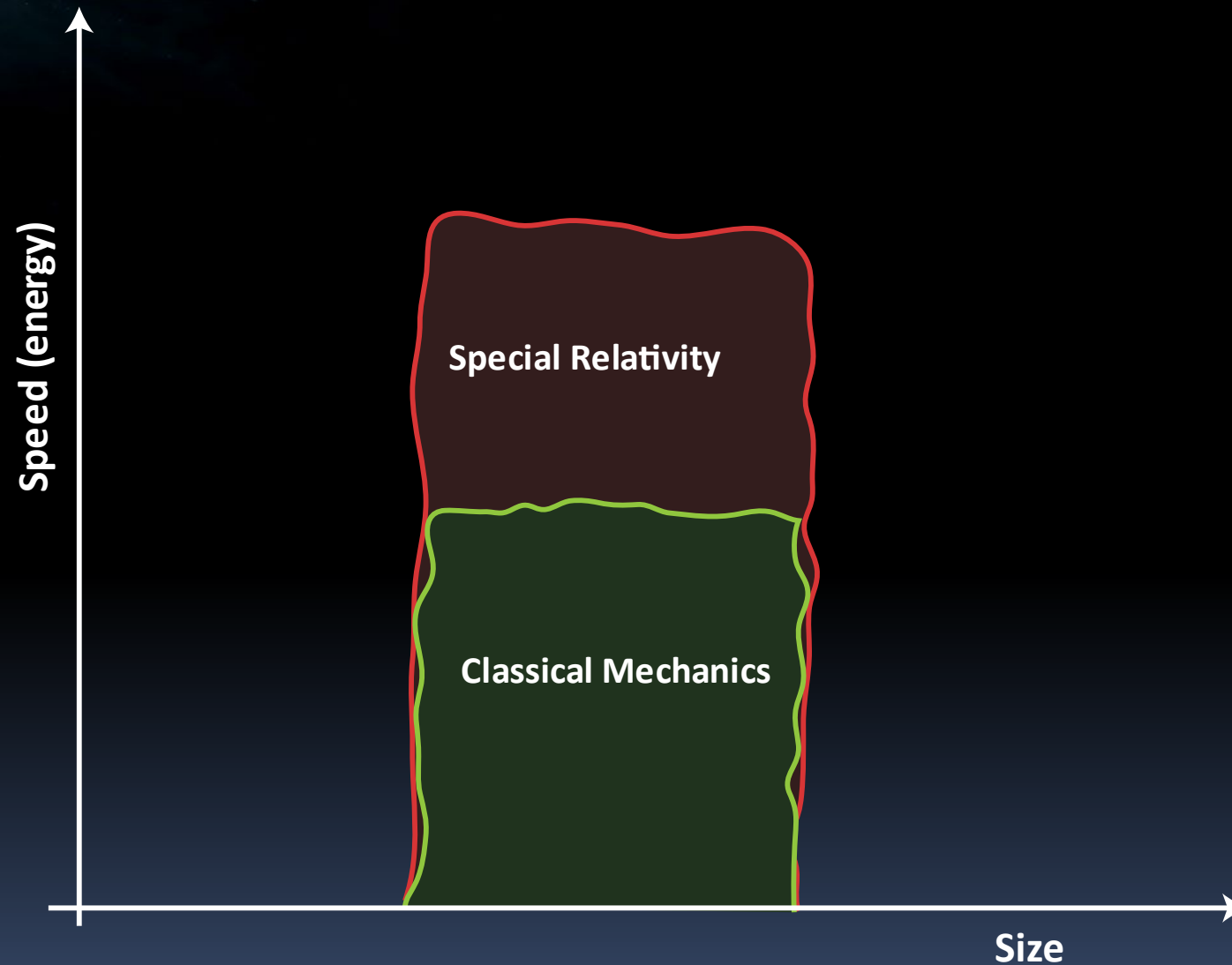
Particles: Small and Fast!



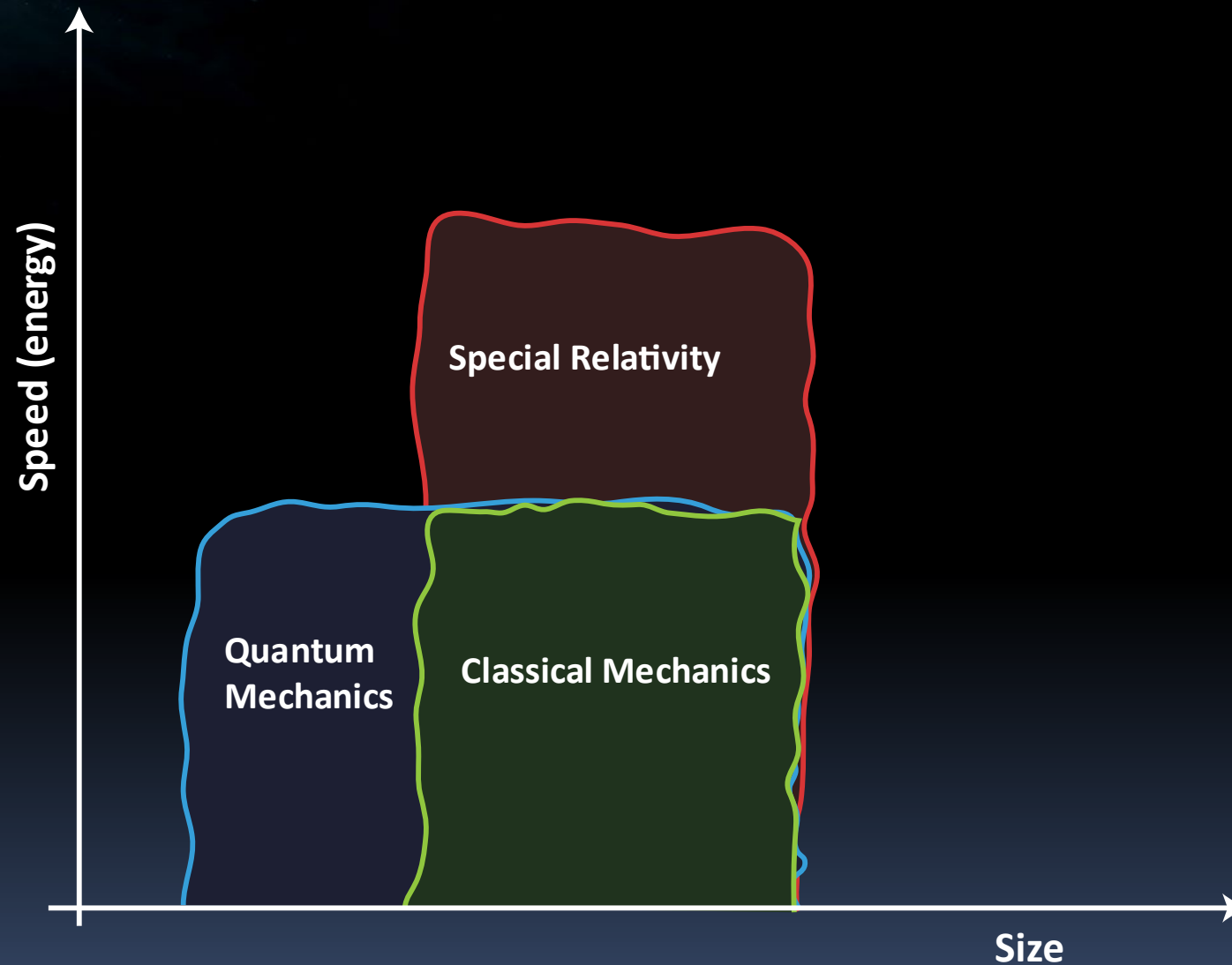
Particles: Small and Fast!



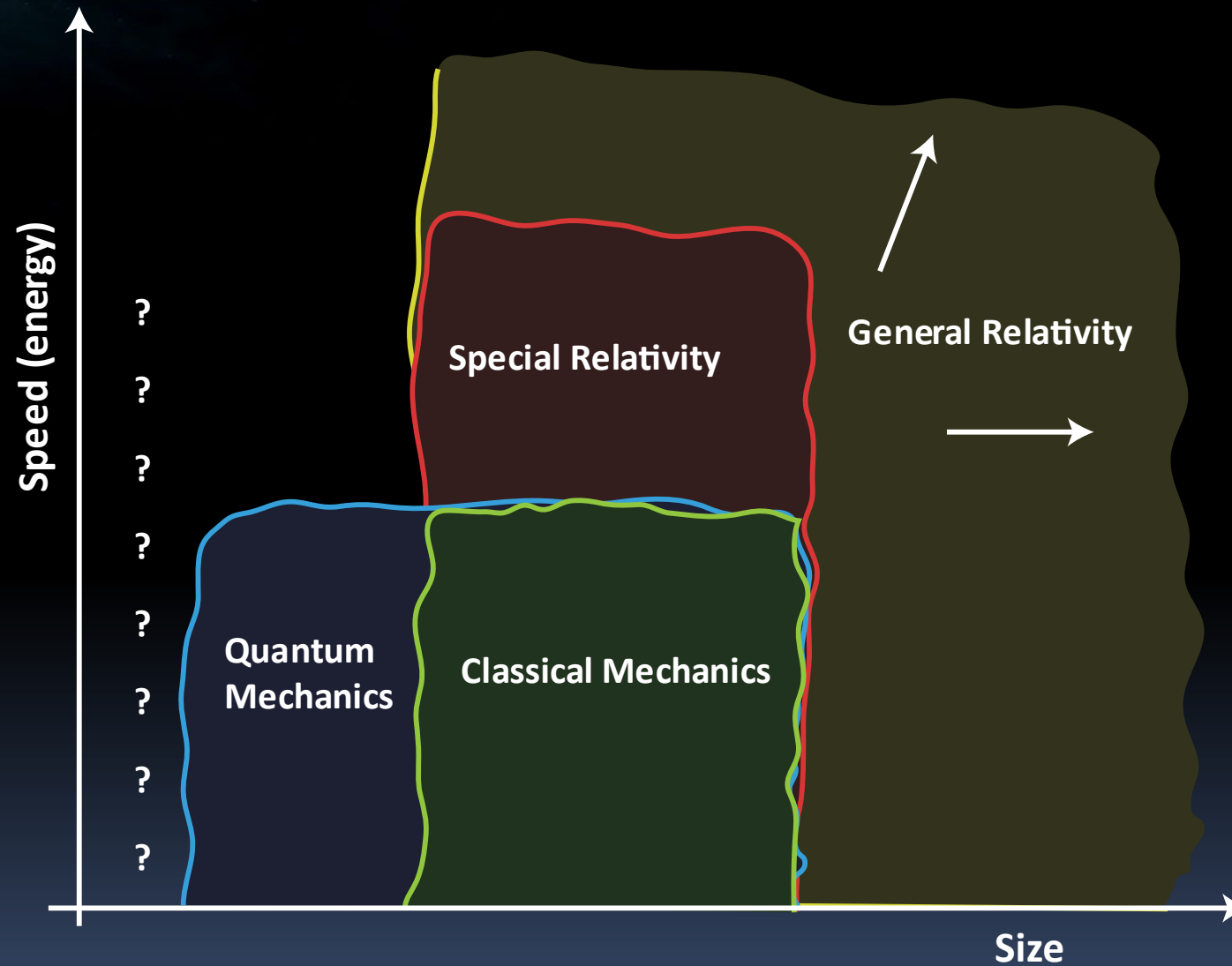
Particles: Small and Fast!



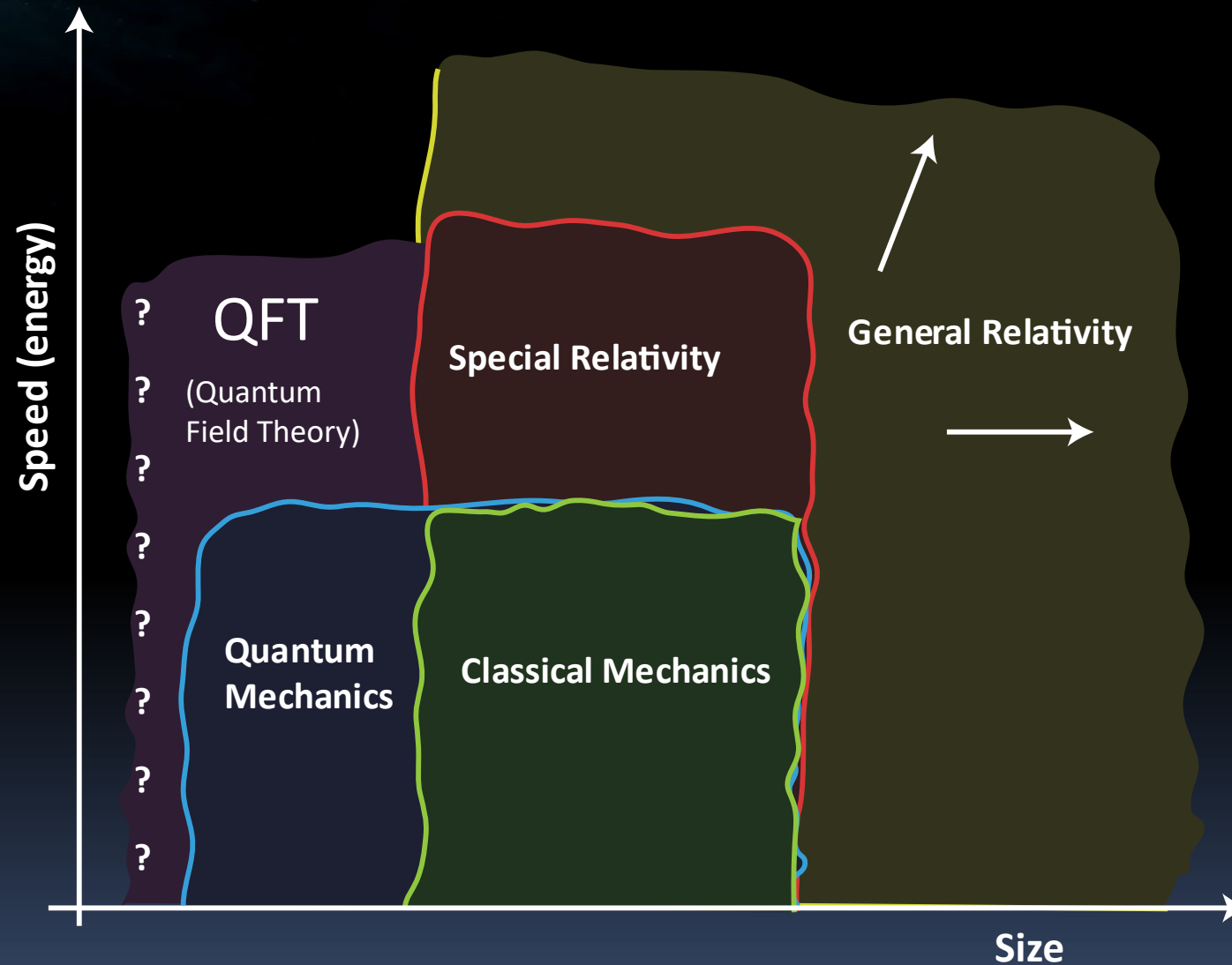
Particles: Small and Fast!



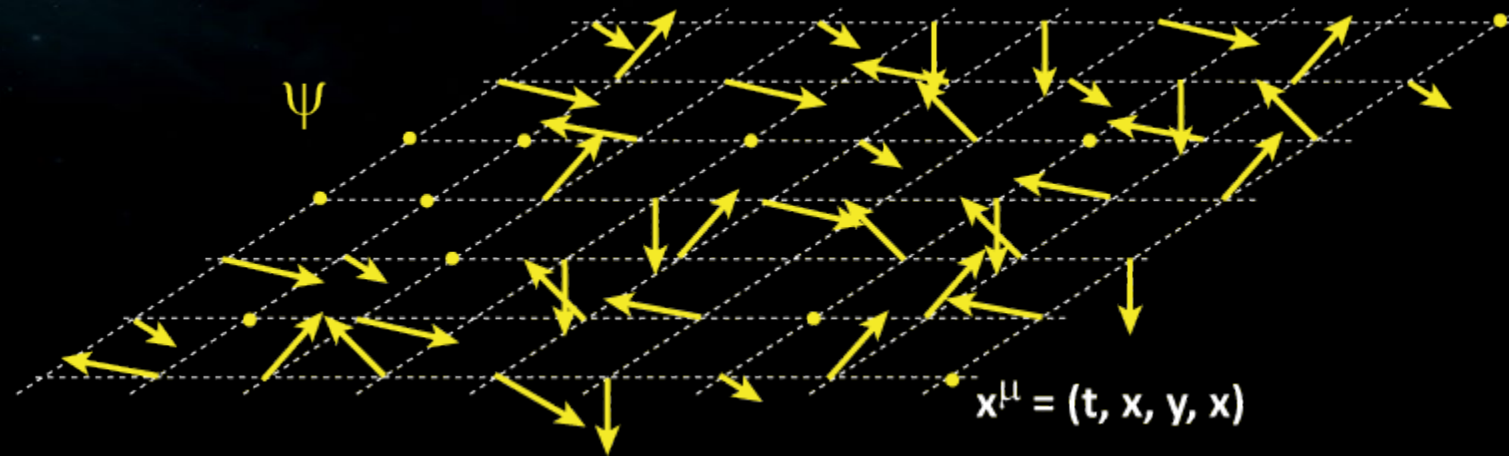
Particles: Small and Fast!



Particles: Small and Fast!



Fields

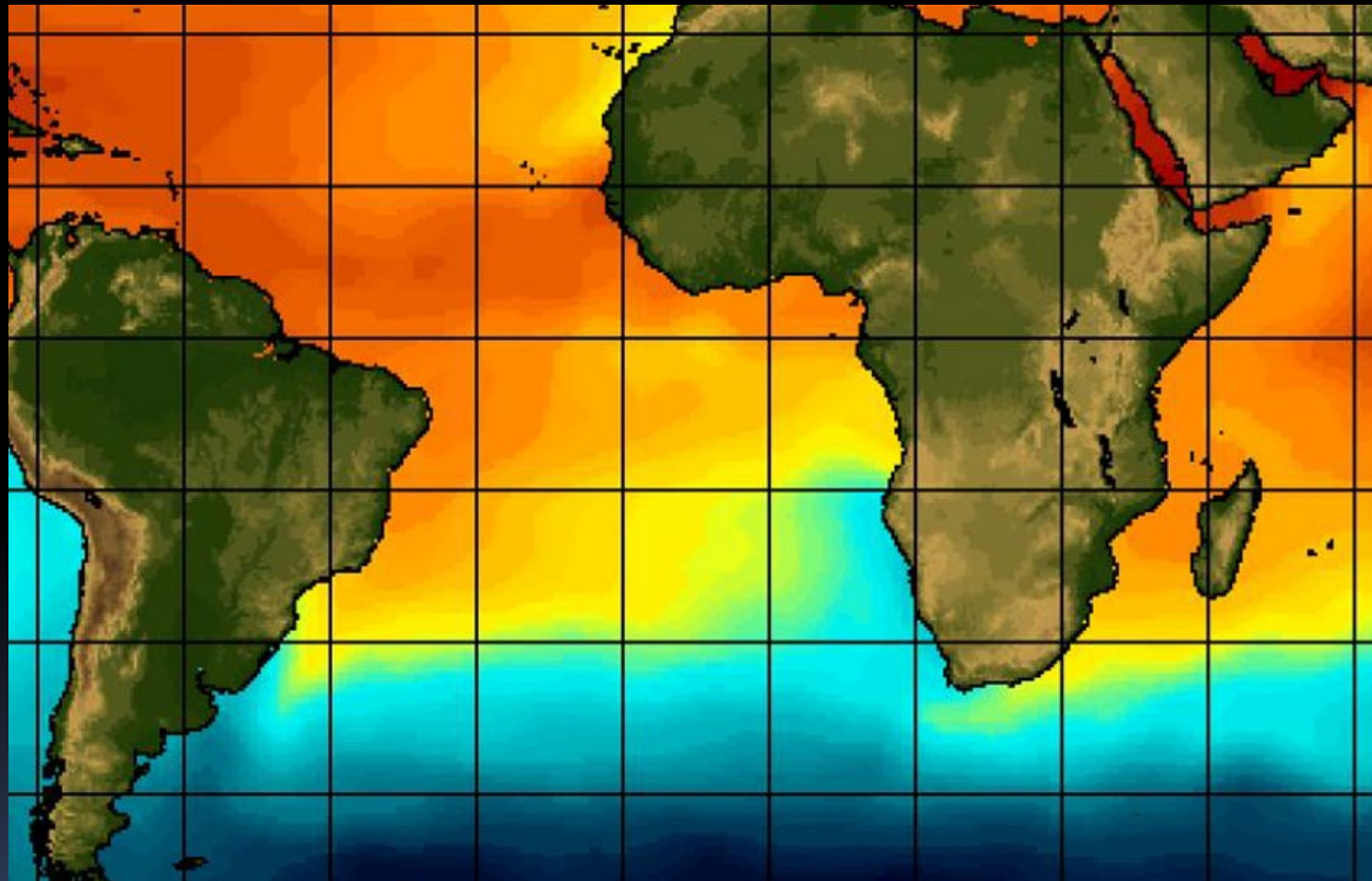


Many types of Fields are possible:

- Scalar. Ex: Temperature, Energy
- Vector. Ex: Electric, Velocities in some fluid or gas
- and more...

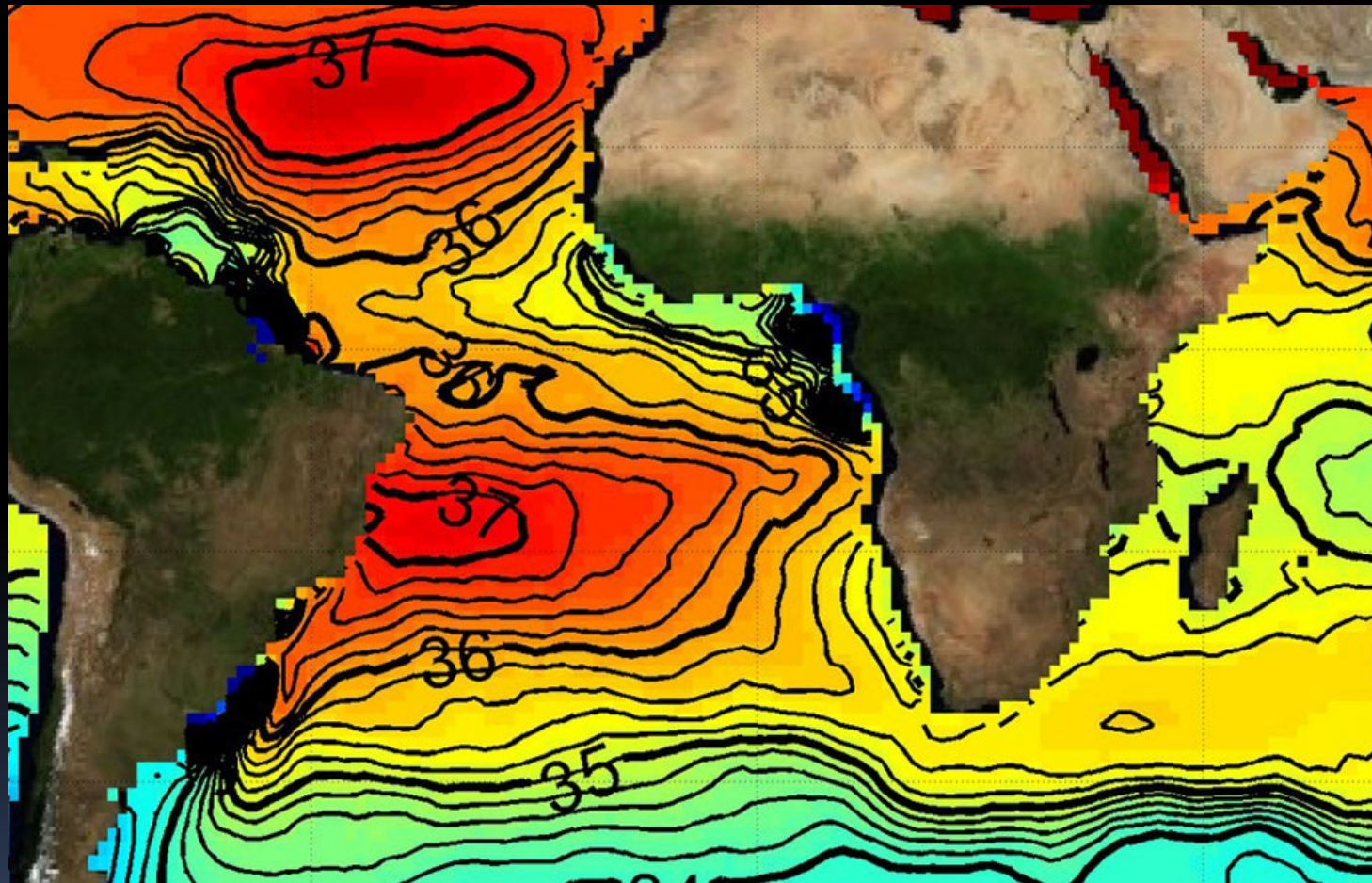
Fields

Example: temperature of oceans (scalar field)



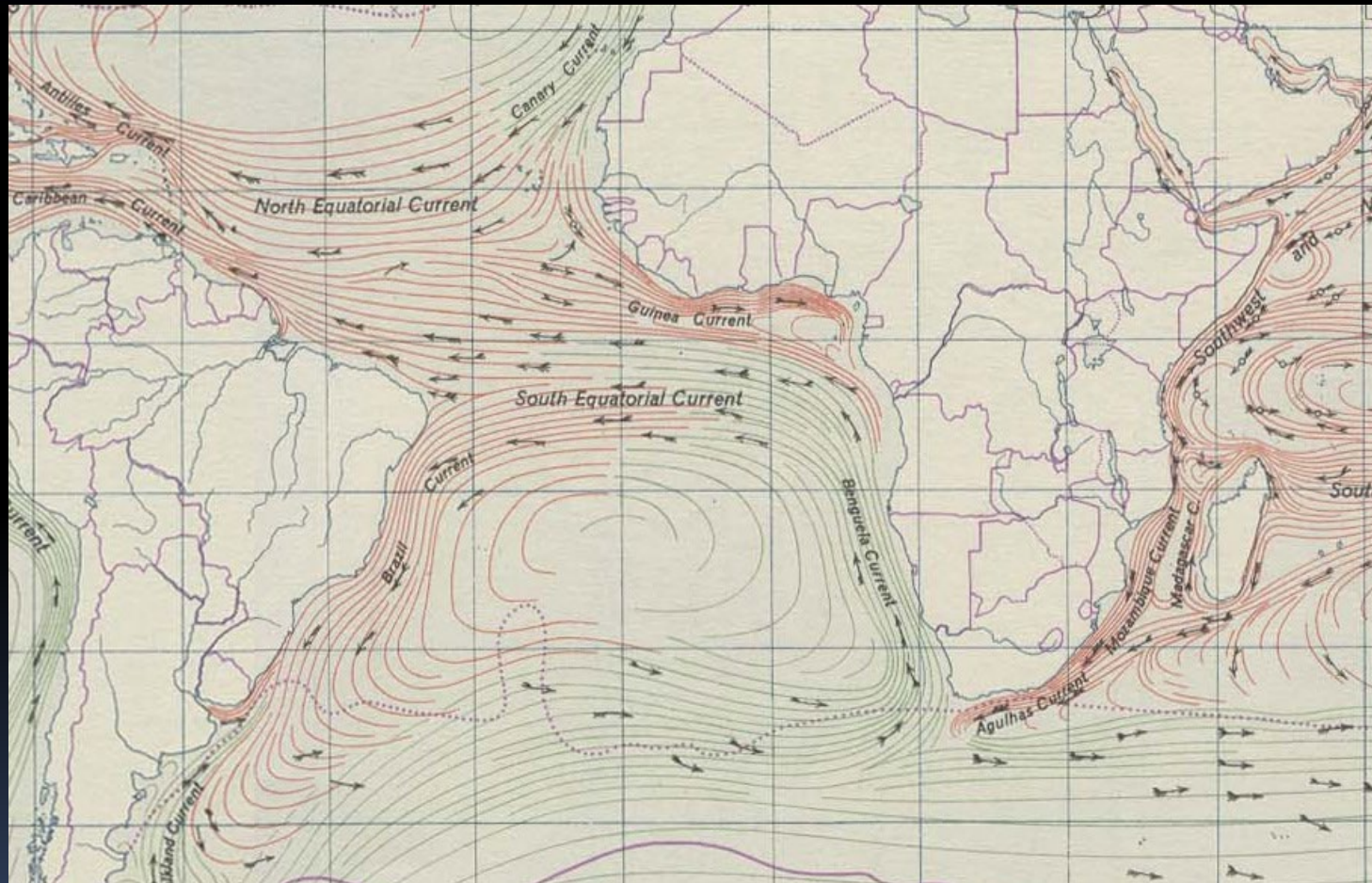
Fields

Example: salinity of oceans (scalar field)



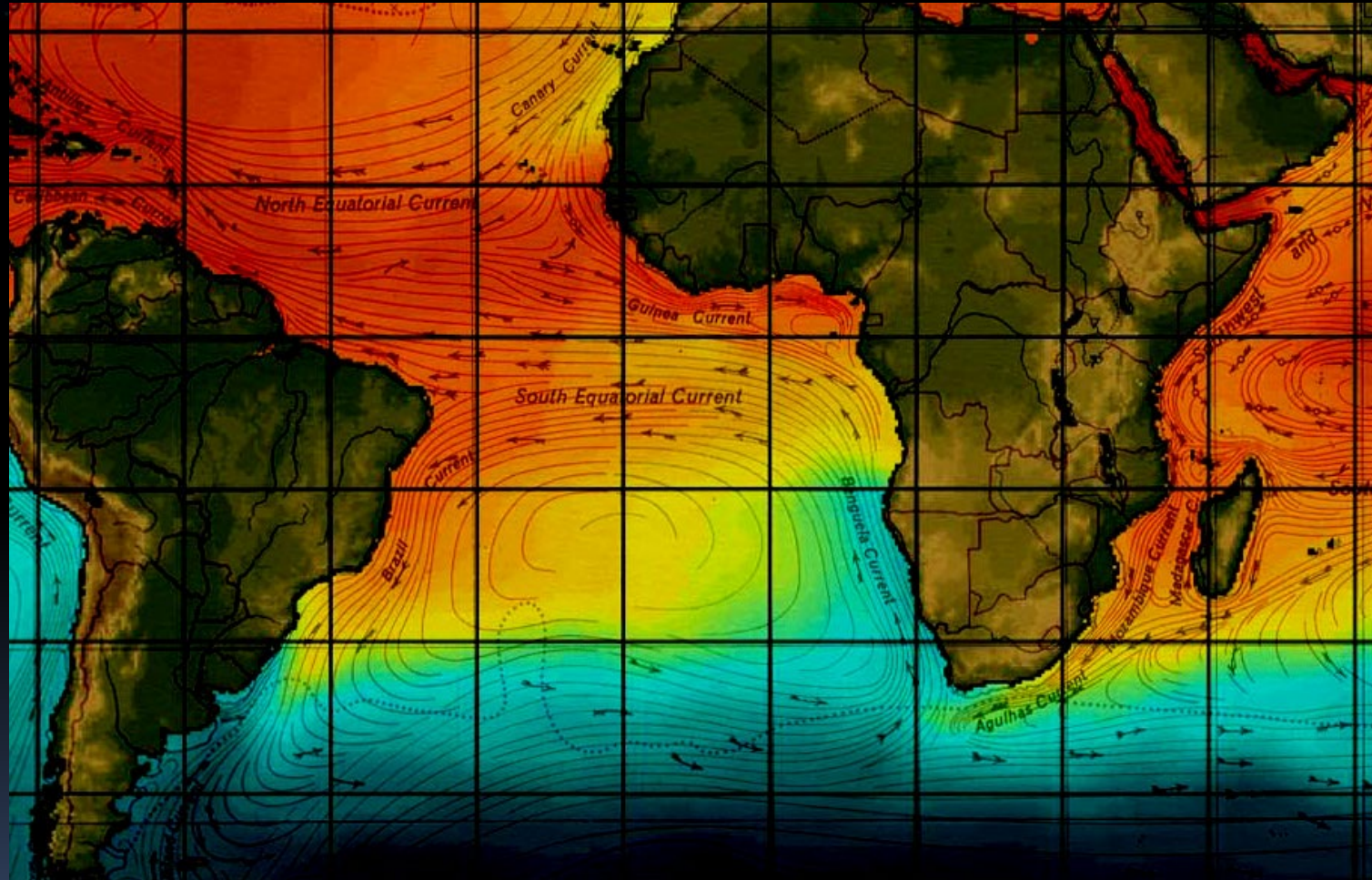
Fields

Example: ocean currents (vector field)

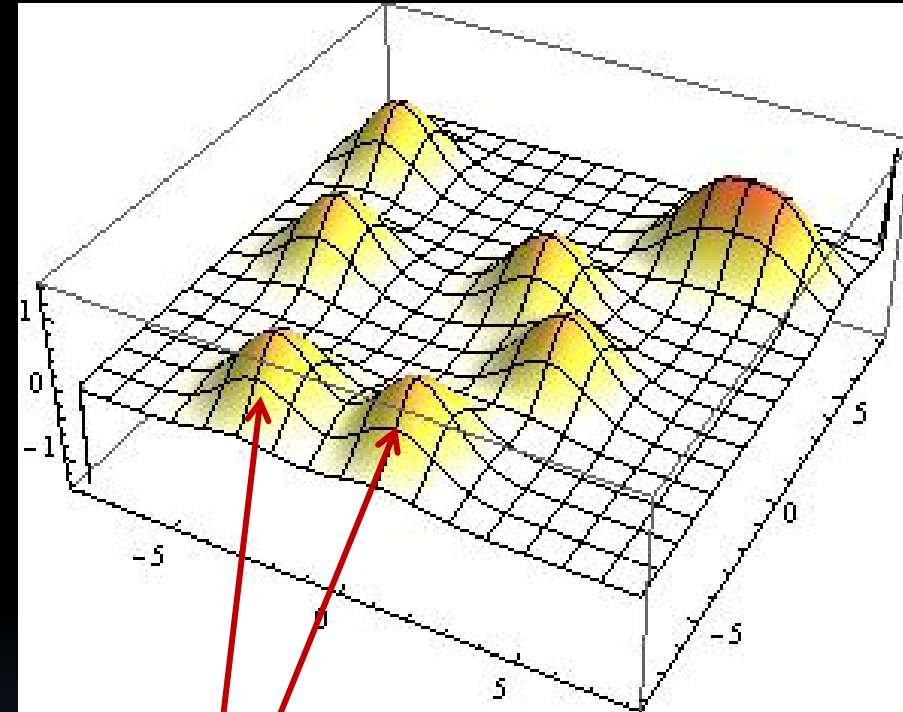
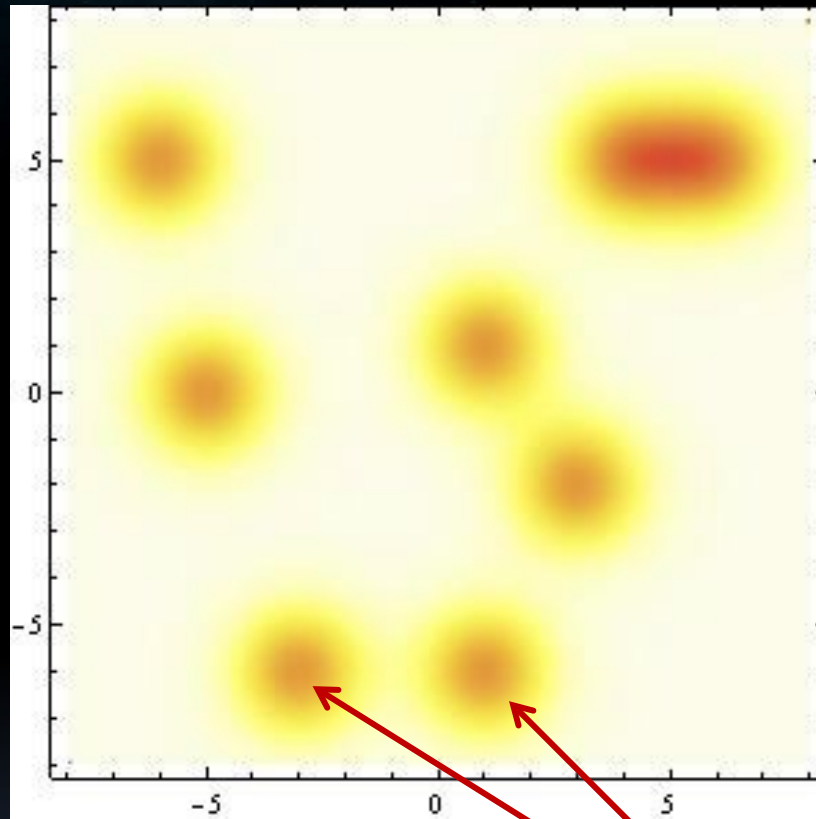


Fields

Can we build a combined field that encapsulates every oceanic property?

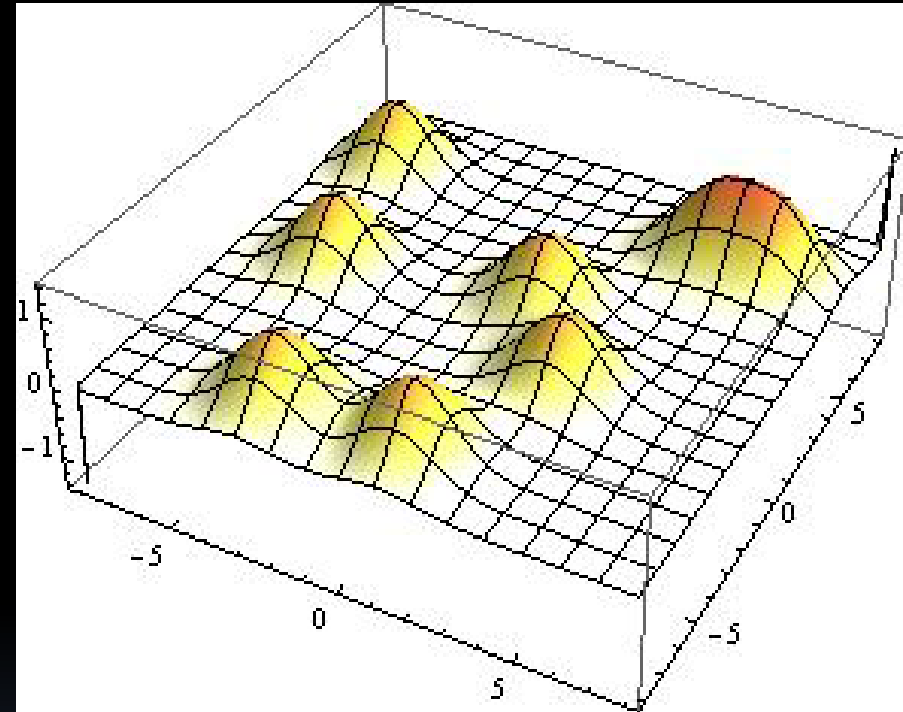
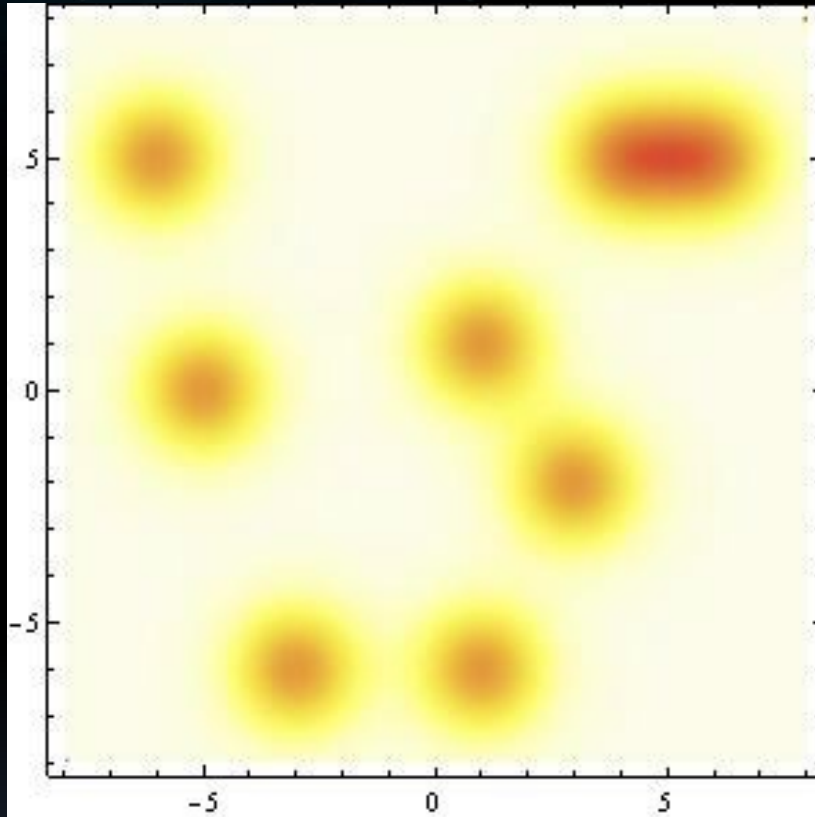


We did it for particles!



Particles

We did it for particles!



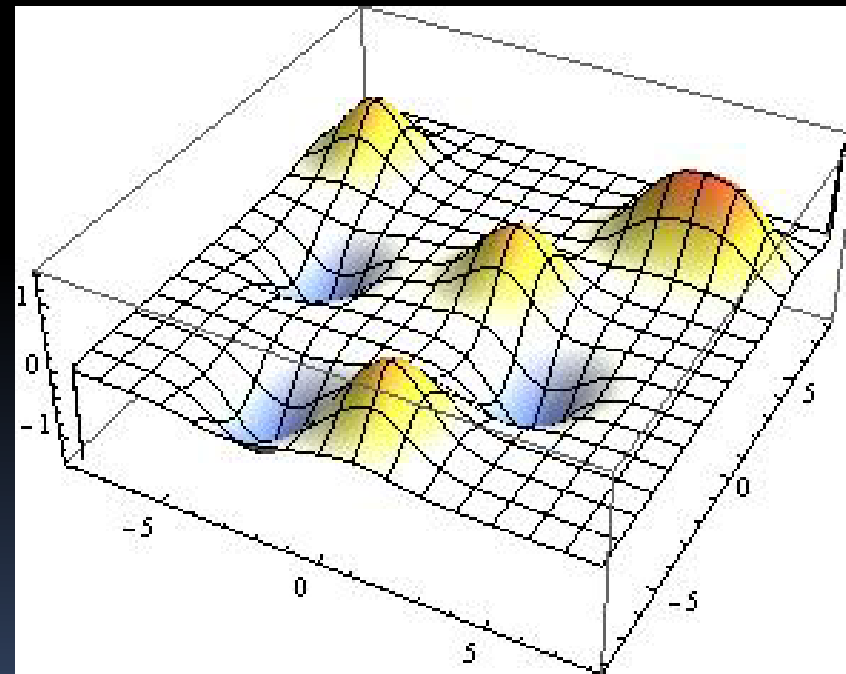
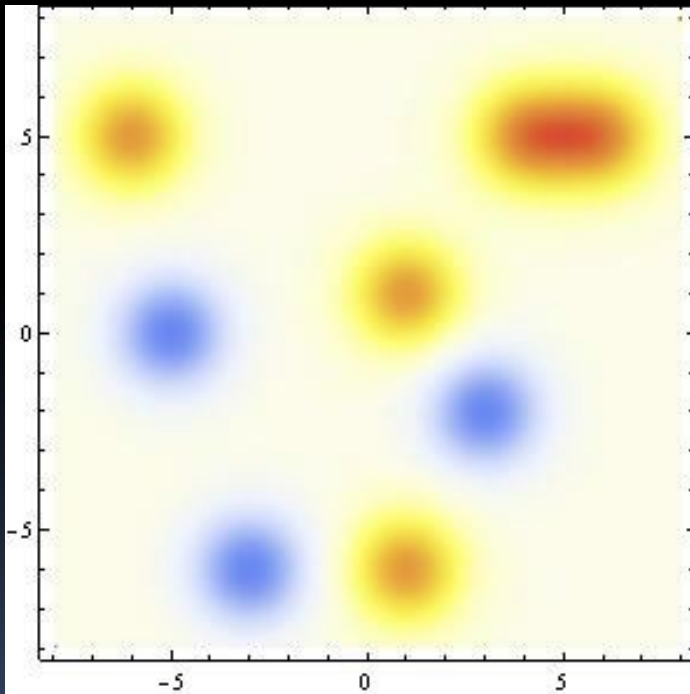
These fields encode not only the probabilities of finding particles in a region of space-time, but also information on velocity, spin, and other quantum numbers.

The simplest consistent Quantum & Relativistic description we know!

Relativistic & Quantum Fields

We get interesting effects both from the quantum side and the relativistic side. A couple of examples:

- Quantized excitations with a relativistic dispersion relation (**particles!**)
- Anti-particles are obligatory! Total number of particles change with time!



Relativistic & Quantum Fields



(note what happens when these excitations cross each other)

Relativistic & Quantum Fields

We get interesting effects both from the quantum side and the relativistic side. A couple of examples:

- Complex Fields

$Z = a + b i = r e^{i \theta}$

The diagram illustrates the conversion of a complex number Z from its rectangular form $a + bi$ to its polar form $re^{i\theta}$. A red arrow points from the underlined text "Complex Fields" to the equation. Another red arrow points from the variable r in the polar form to the word "Module". A third red arrow points from the exponent $i\theta$ to the word "Phase".

Relativistic & Quantum Fields

We get interesting effects both from the quantum side and the relativistic side. A couple of examples:

- Complex Fields

$$Z = a + b i = r e^{i\theta}$$

Phase

Module

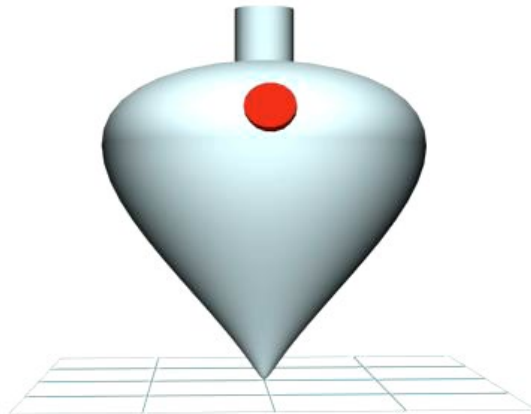
However, our observations are REAL, and do not depend on this phase

↓
Symmetry!!



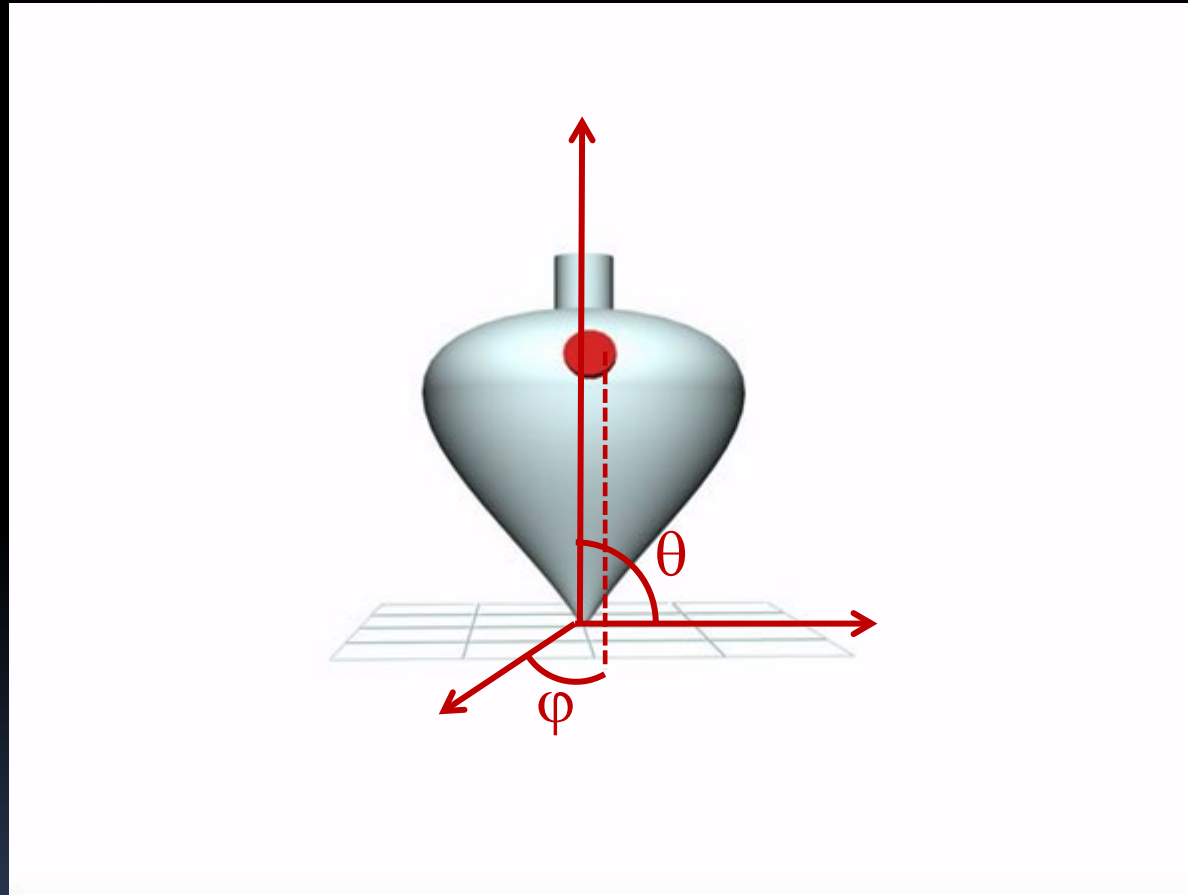
Role of Symmetries in Physics

What is a symmetry?



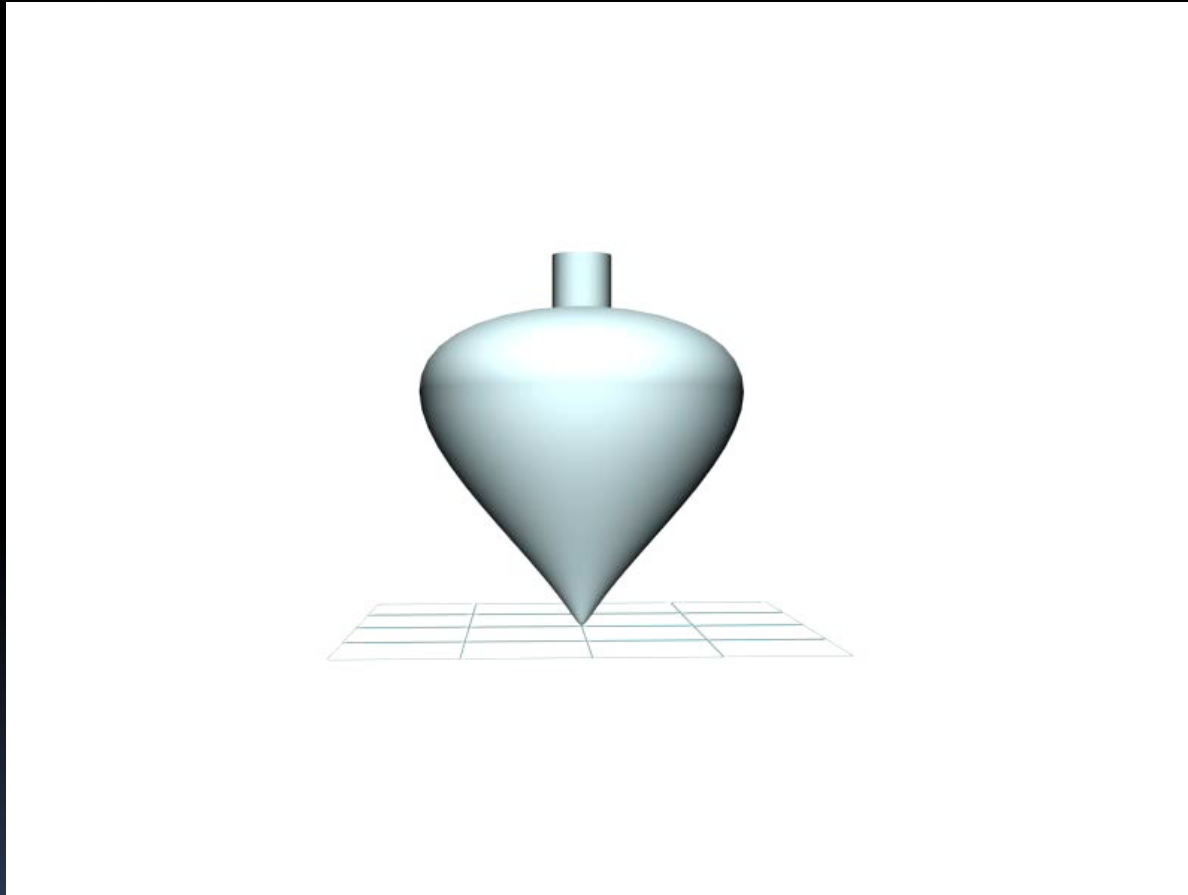
Role of Symmetries in Physics

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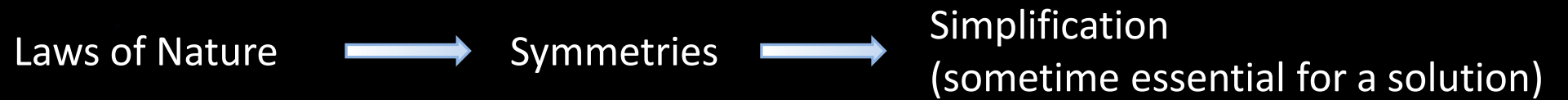
Role of Symmetries in Physics

What is a symmetry?



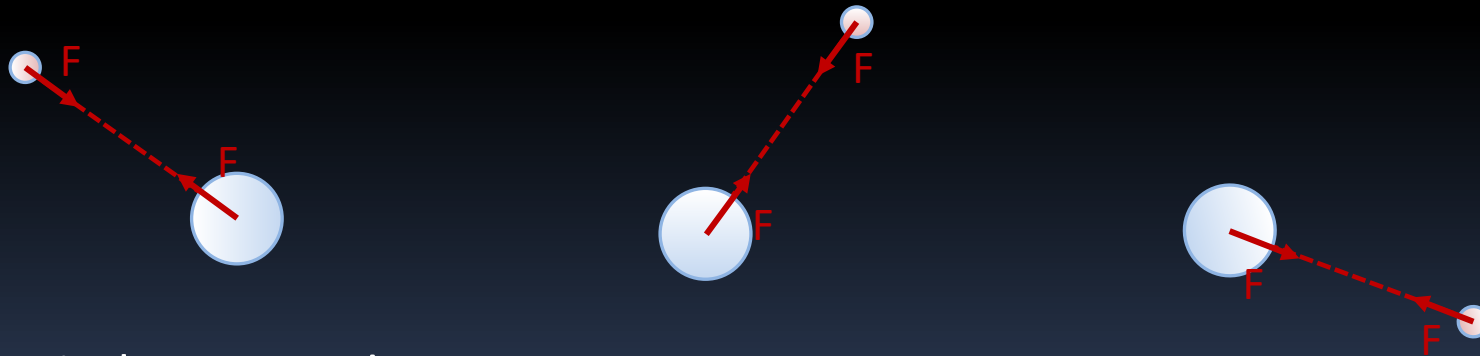
Role of Symmetries in Physics

Before the XX century:



Example: Newton's Gravitation

...force is proportional to the masses and inversely proportional to the square of the distance between bodies

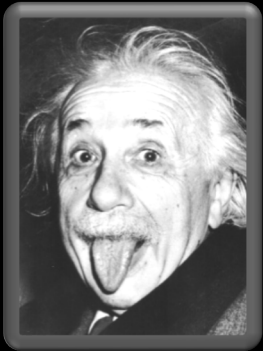


Spherical Symmetry!


Role of Symmetries in Physics

Start of XX century: roles start to change

Symmetry  Laws of Nature



A. Einstein – 1905 ~ 1915

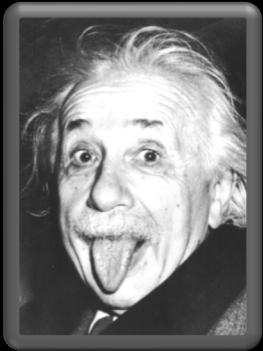
 Relativity (Special and General) is derived for symmetry principles
(laws should be reference frame independent)

 Restricts possible “laws”


Role of Symmetries in Physics

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Symmetry  Laws of Nature



A. Einstein – 1905 ~ 1915

 Relativity (Special and General) is derived for symmetry principles (laws should be reference frame independent)

 Restricts possible “laws”

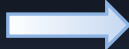
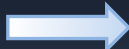



A.E. Noether- 1918

 Symmetries imply **conservation law**



Invariance under:

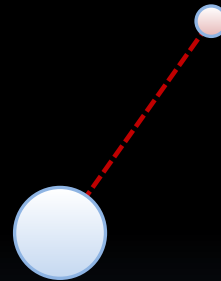
Translations in space		Momentum conservation
Translations in time		Energy conservation
Rotations in 3D		Angular momentum conservation

Role of Symmetries in Physics

Start of XX century: roles start to change

Symmetry  Laws of Nature

Example: Newton's Gravitation



Role of Symmetries in Physics

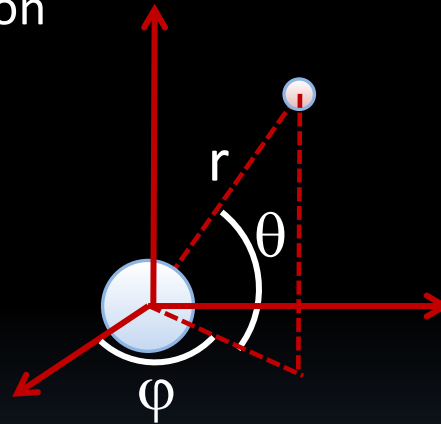
Start of XX century: roles start to change

Symmetry



Laws of Nature

Example: Newton's Gravitation

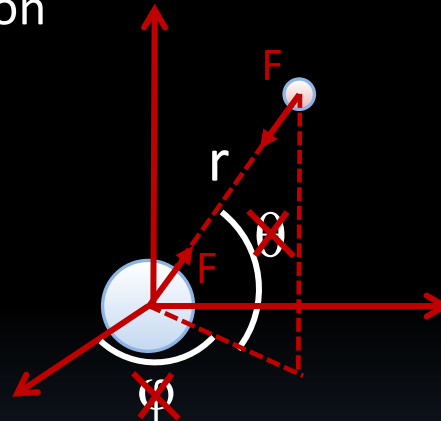


Role of Symmetries in Physics

Start of XX century: roles start to change

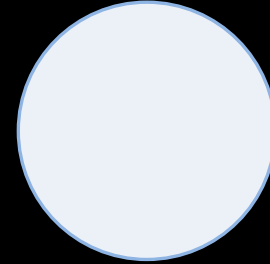
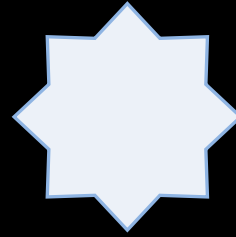
Symmetry \longrightarrow Laws of Nature

Example: Newton's Gravitation

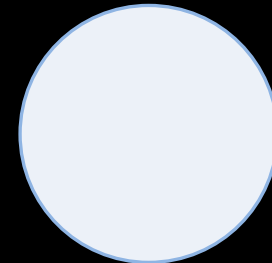
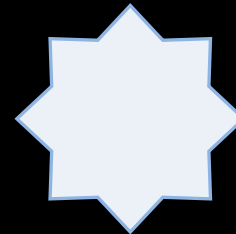
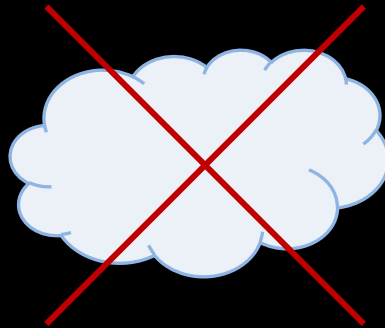
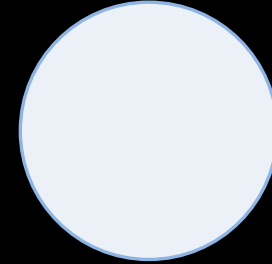
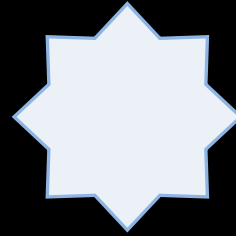


Spherical Symmetry \longrightarrow Force depends only on distance

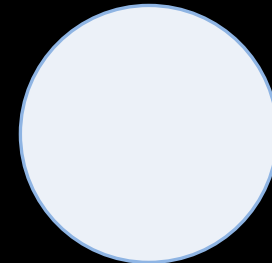
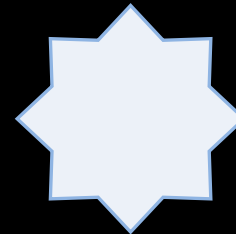
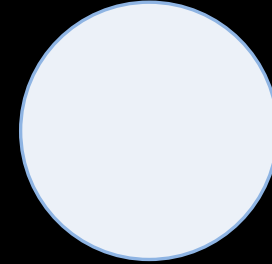
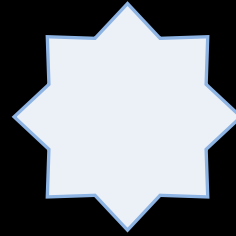
Symmetries → Restrictions



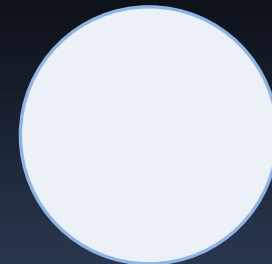
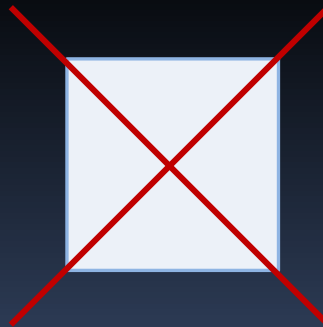
Symmetries Restrictions



Symmetries Restrictions



90°



45°

Role of Symmetries in Physics



Complex Field:

$$Z = a + b i = r e^{i \theta}$$

→ Theory must be
symmetric **phase**



Noether: Symmetry → Conserved Quantity

Role of Symmetries in Physics



Complex Field:

$$Z = a + b i = r e^{i \theta}$$

Theory must be
symmetric **phase**



Noether: Symmetry \longrightarrow Conserved Quantity

Conservation of **ELECTRIC CHARGE**

Role of Symmetries in Physics



Complex Field:

$$Z = a + b i = r e^{i \theta}$$

Theory must be
symmetric **phase**



Noether: Symmetry \longrightarrow Conserved Quantity

Conservation of **ELECTRIC CHARGE**

Complex fields describe charged particles!



(the electron field also describes **spin 1/2** particles, fields of this type are called **spinor fields**, and are different from vector or scalar fields)

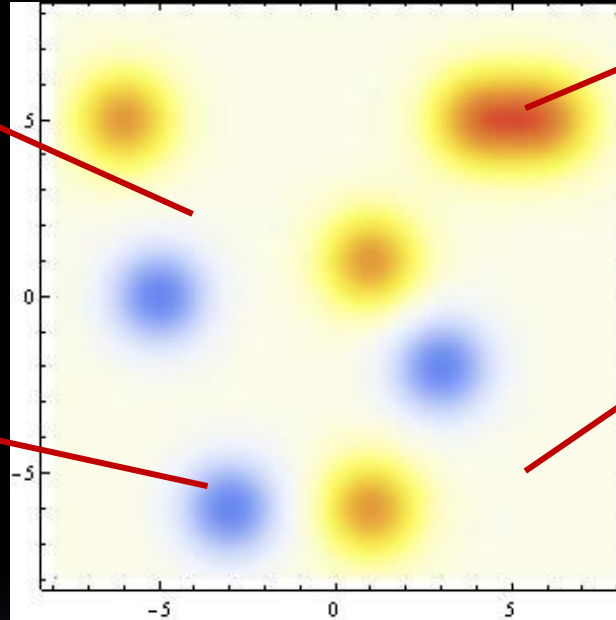
Global vs Local Symmetries

$$Z_1 = r_1 e^{i(\theta_1)}$$

$$Z_2 = r_2 e^{i(\theta_2)}$$

$$Z_3 = r_3 e^{i(\theta_3)}$$

$$Z_4 = r_4 e^{i(\theta_4)}$$



Global Transformation

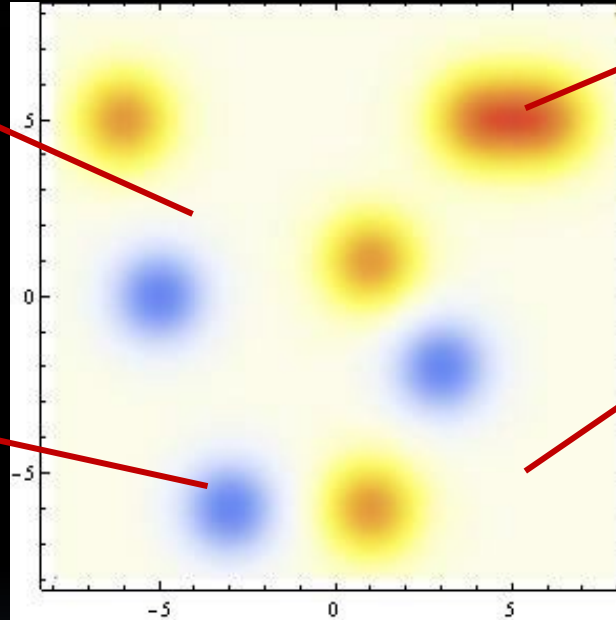
Global vs Local Symmetries

$$Z_1 = r_1 e^{i(\theta_1 + \alpha)}$$

$$Z_2 = r_2 e^{i(\theta_2 + \alpha)}$$

$$Z_3 = r_3 e^{i(\theta_3 + \alpha)}$$

$$Z_4 = r_4 e^{i(\theta_4 + \alpha)}$$



Global Transformation

Conservation of **ELECTRIC CHARGE**

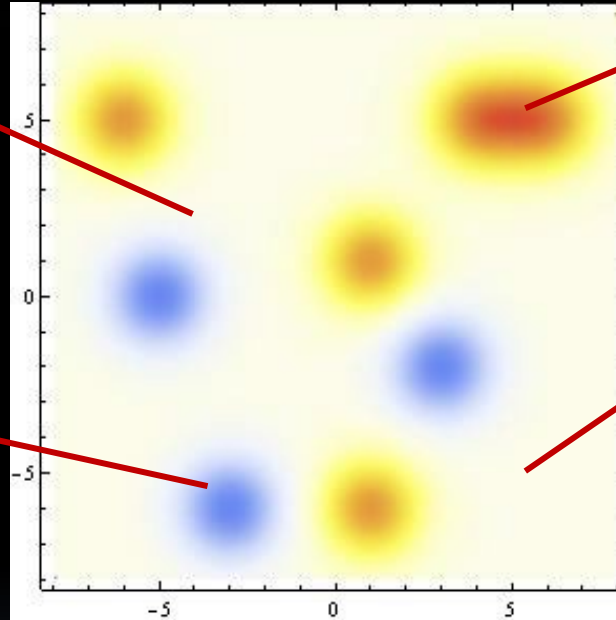
Global vs Local Symmetries

$$Z_1 = r_1 e^{i(\theta_1 + \alpha)}$$

$$Z_2 = r_2 e^{i(\theta_2 + \beta)}$$

$$Z_3 = r_3 e^{i(\theta_3 + \gamma)}$$

$$Z_4 = r_4 e^{i(\theta_4 + \delta)}$$



Local Transformation

???

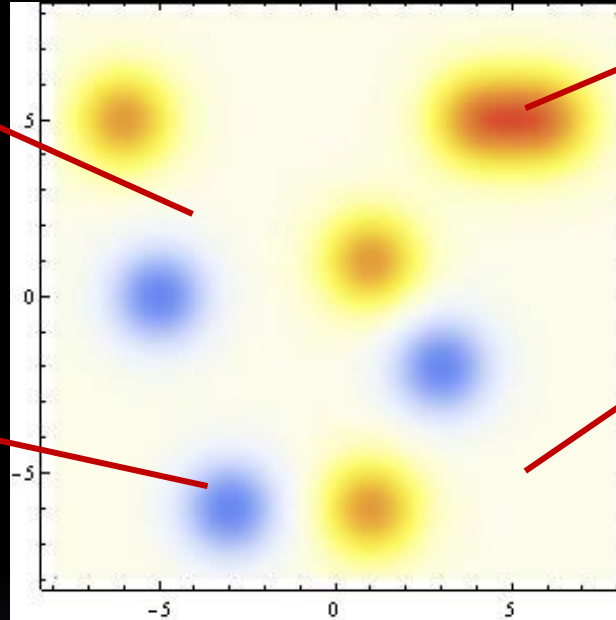
Global vs Local Symmetries

$$Z_1 = r_1 e^{i(\theta_1 + \alpha)}$$

$$Z_2 = r_2 e^{i(\theta_2 + \beta)}$$

$$Z_3 = r_3 e^{i(\theta_3 + \gamma)}$$

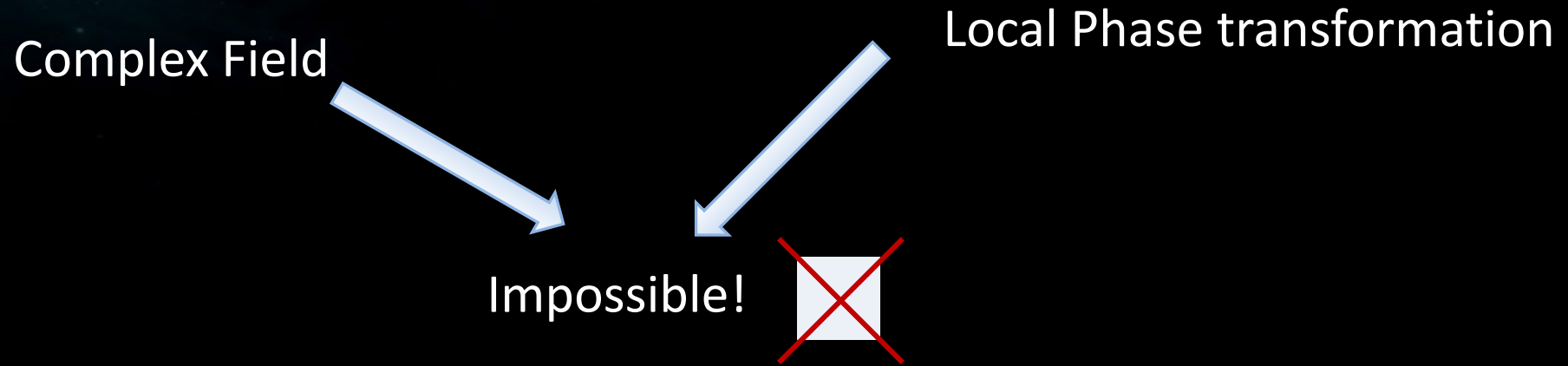
$$Z_4 = r_4 e^{i(\theta_4 + \delta)}$$



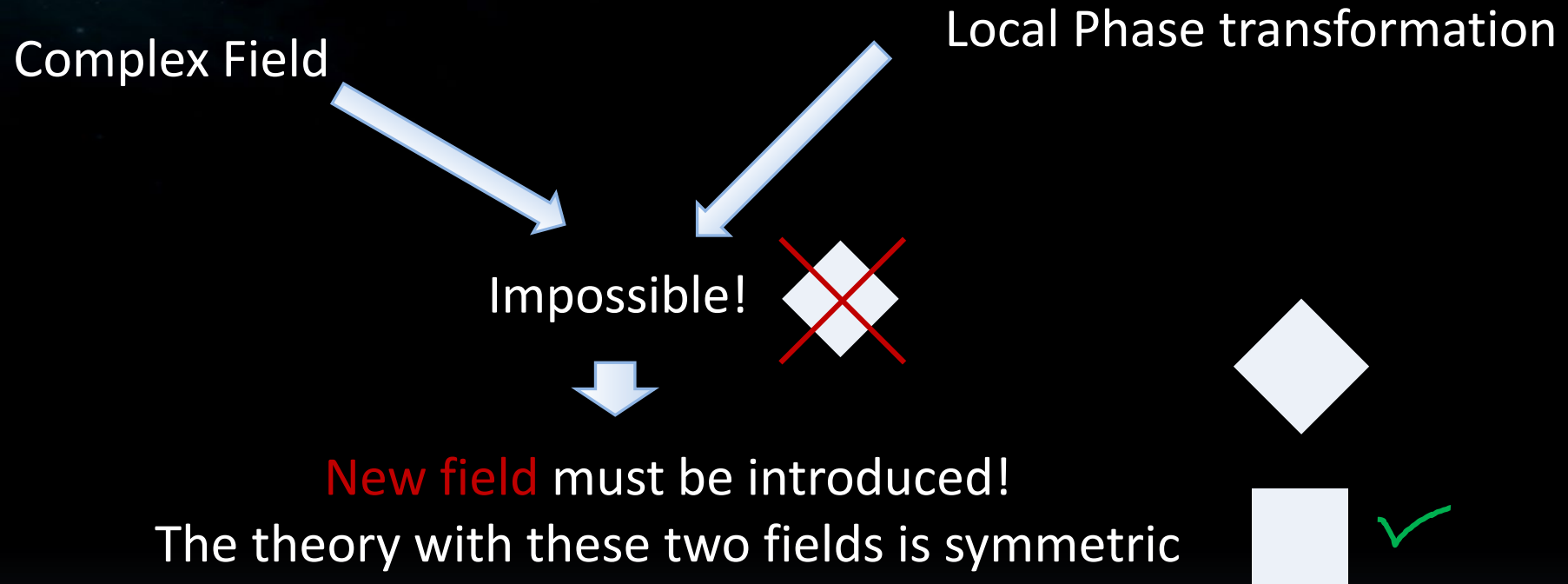
Local Transformation

INTERACTIONS!!

Local Symmetries and Gauge Bosons



Local Symmetries and Gauge Bosons



Local Symmetries and Gauge Bosons

Complex Field

Local Phase transformation


Impossible!



New field must be introduced!

The theory with these two fields is symmetric



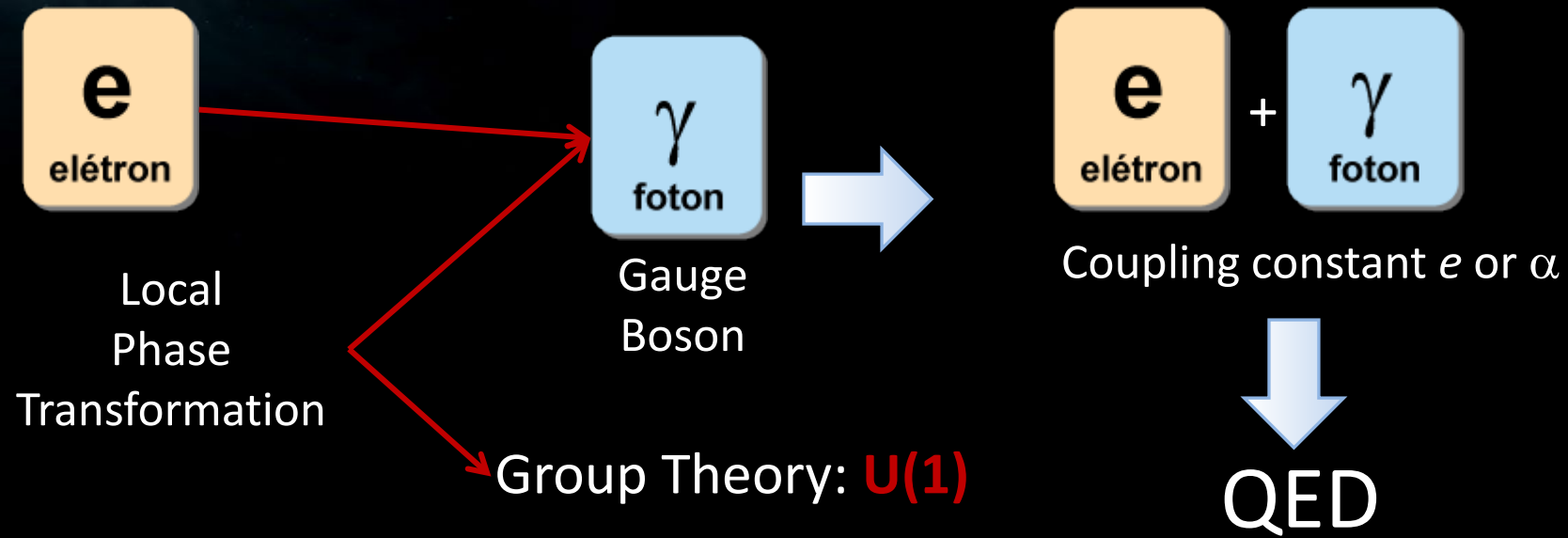
- Vector Field
- Field excitations **HAVE NO MASS** 
- Interaction is closely related to the **conserved charge and gauge coupling**

Local Symmetries and Gauge Bosons

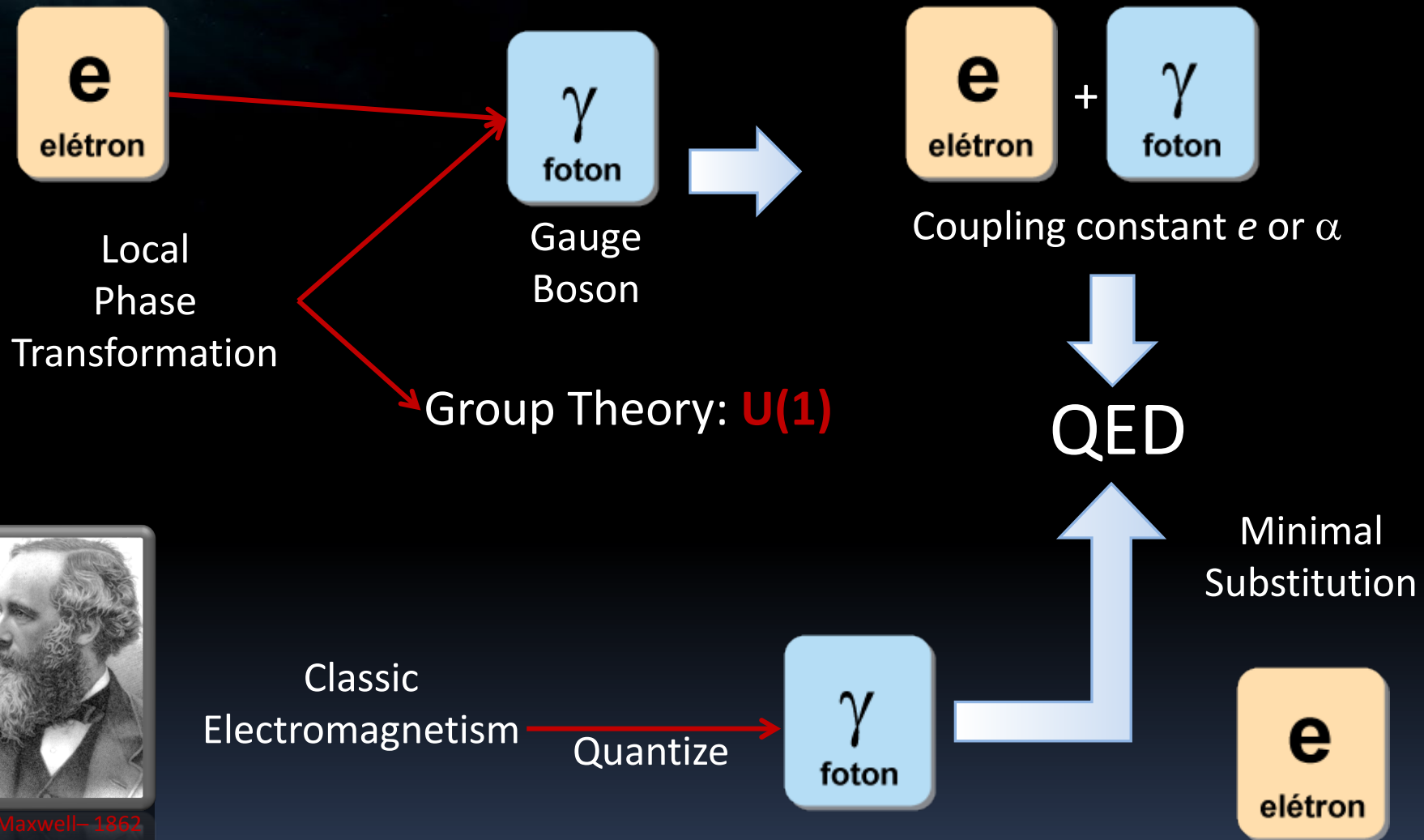


Vocabulary: Gauge Symmetries, Gauge Fields and Gauge Couplings

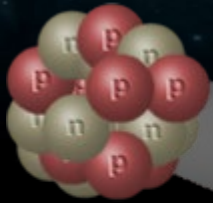
Quantum Electrodynamics (QED)



Quantum Electrodynamics (QED)



Quantum Chromodynamics (QCD)



Quarks

What holds the quarks together?

1963



G. Zweig



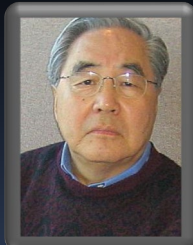
M. Gell-Mann

Local Symmetry: **SU(3)**

1965



Y. Nambu



M-Y Han



O.W. Greenberg

SU(3) charged

u
up

d
down

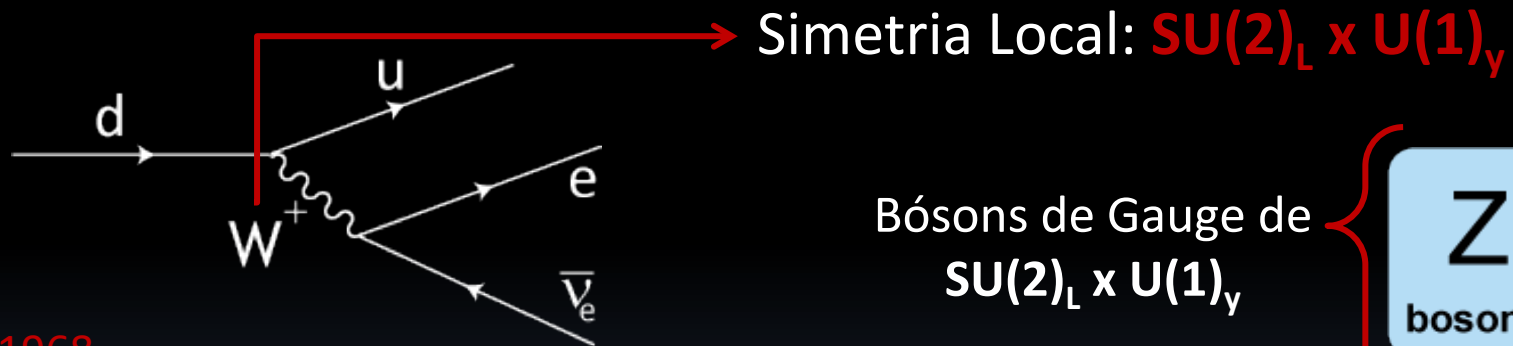
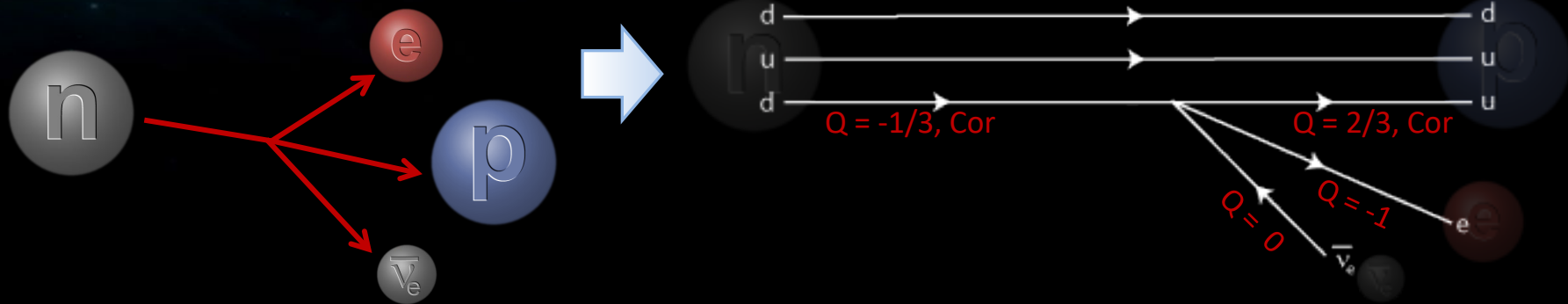
s
strange

SU(3)
Gauge Boson

g
gluon

Electroweak Model (GWS)

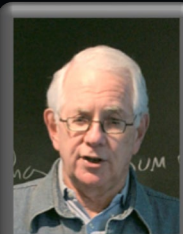
Decaimento Beta:



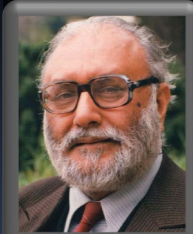
Bósons de Gauge de $SU(2)_L \times U(1)_Y$



1968



Glashow



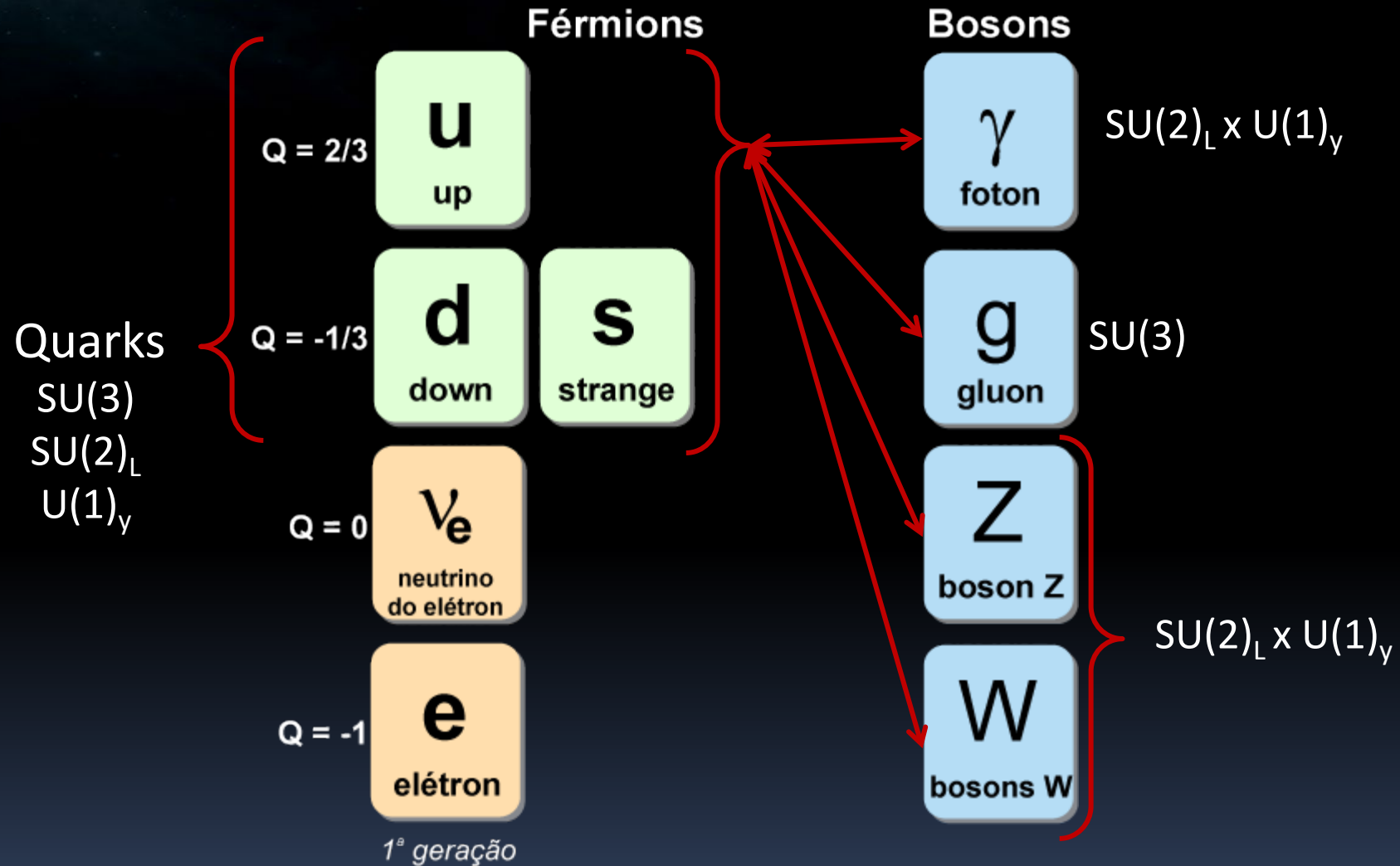
Salan



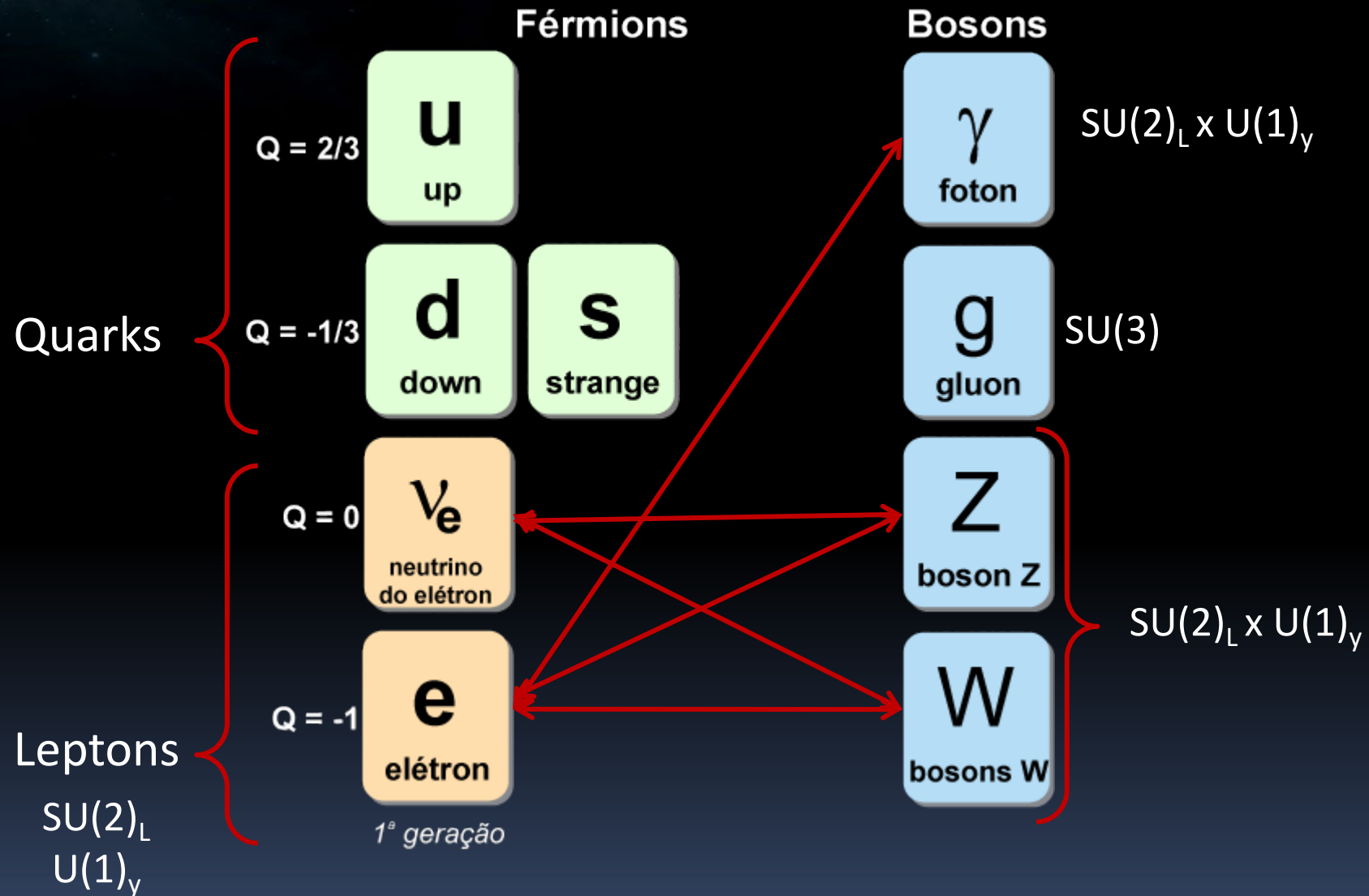
Weinberg

A MODEL OF LEPTONS*
Steven Weinberg†
Laboratory for Nuclear Science and Physics Department,
Massachusetts Institute of Technology, Cambridge, Massachusetts
(Received 17 October 1967)

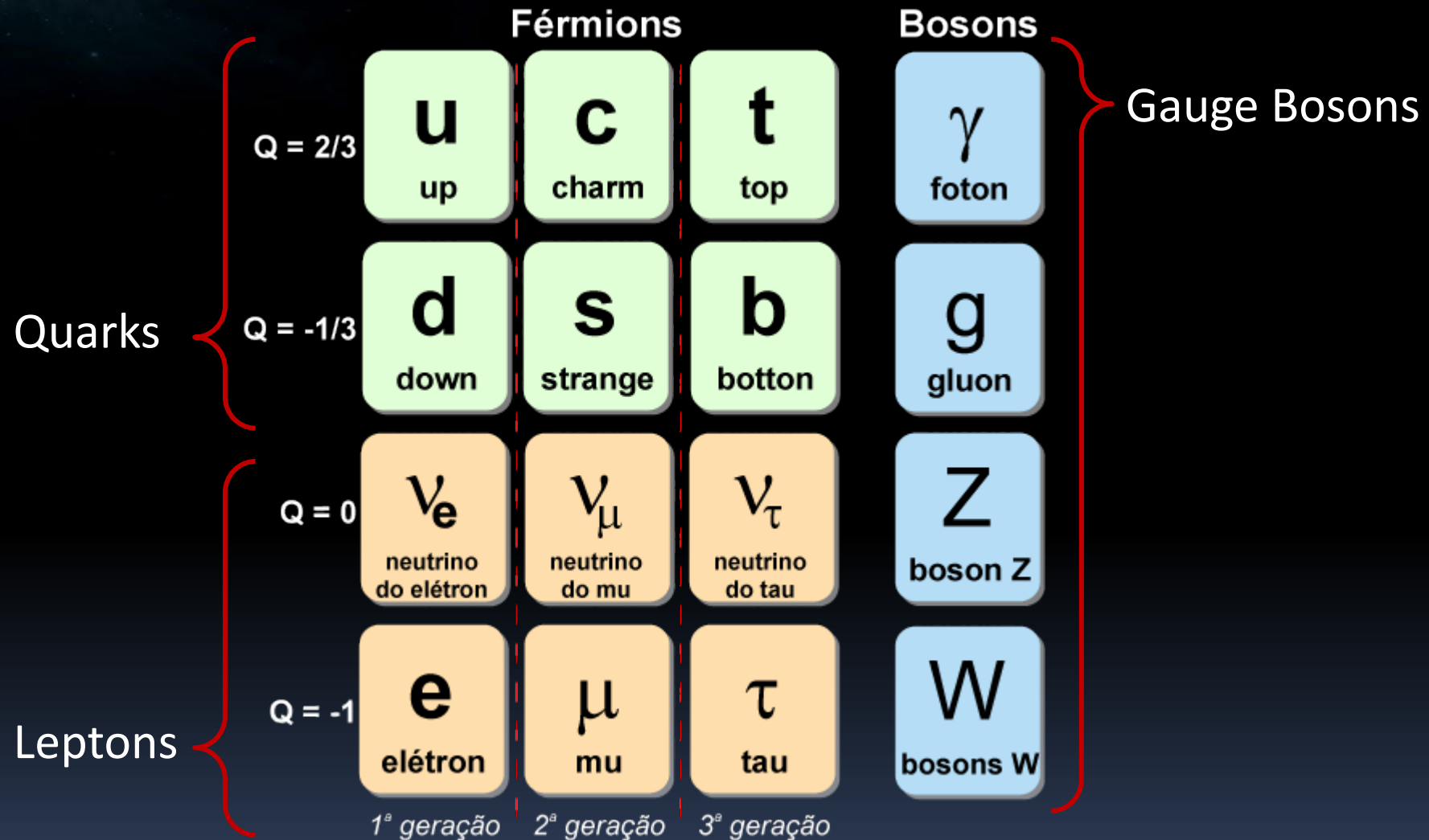
The Standard Model (SM)



The Standard Model (SM)



The Standard Model (SM)



The Standard Model (SM)

Is that all? Can we go home?

Well, time to recall some facts:

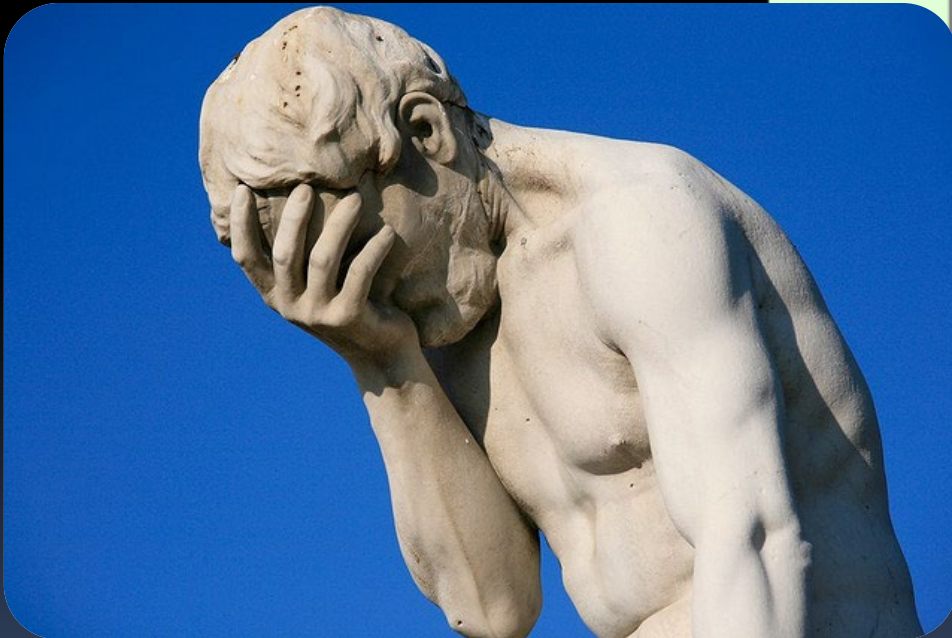


“Gauge Bosons **HAVE NO MASS**”

... let's take a look.

The Standard Model (SM)

	Férmions			Bosons	
$Q = 2/3$	u up	c charm	t top	γ foton	$m = 0$ ✓
$Q = -1/3$	d	s strange	b bottom	g gluon	$m = 0$ ✓
	ν_μ neutrino do mu	ν_τ neutrino do tau		Z boson Z	$m = 92 \text{ GeV}$ ✗
	μ mu	τ tau		W bosons W	$m = 80 \text{ GeV}$ ✗
	2ª geração		3ª geração		



The Standard Model (SM)

Is that all? Can we go home?

Well, time to recall some facts:



“Gauge Bosons **HAVE NO MASS**”

... let's take a look.

W and Z bosons have all the symmetries predicted by $SU(2)_L \times U(1)_Y$, but they have **MASS**

The Standard Model (SM)

Is that all? Can we go home?

Well, time to recall some facts:



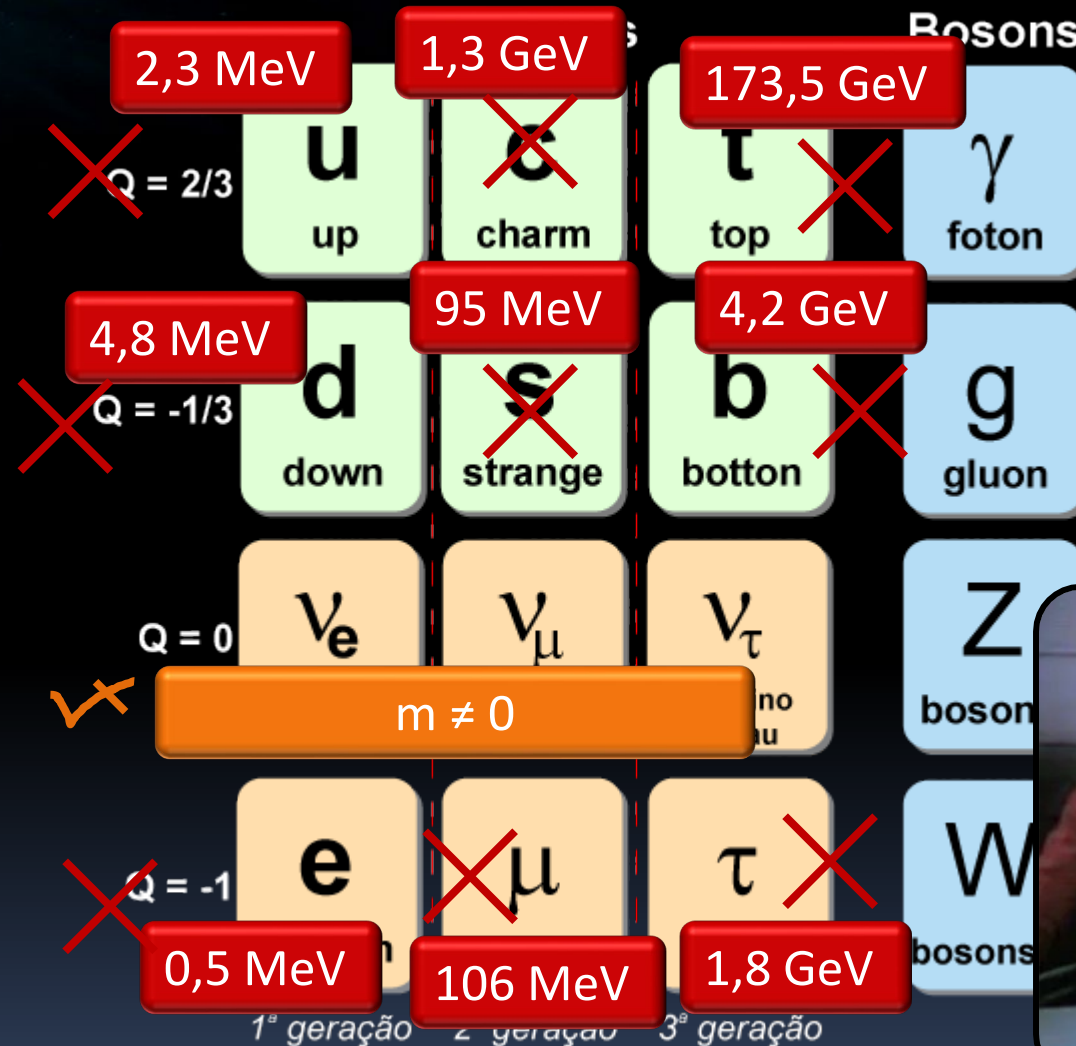
“Gauge Bosons **HAVE NO MASS**”

... let's take a look.

W and Z bosons have all the symmetries predicted by $SU(2)_L \times U(1)_Y$, but they have **MASS**

That is not all: the $SU(2)_L$ symmetry forbids **fermion** masses too! Let's see...

The Standard Model (SM)



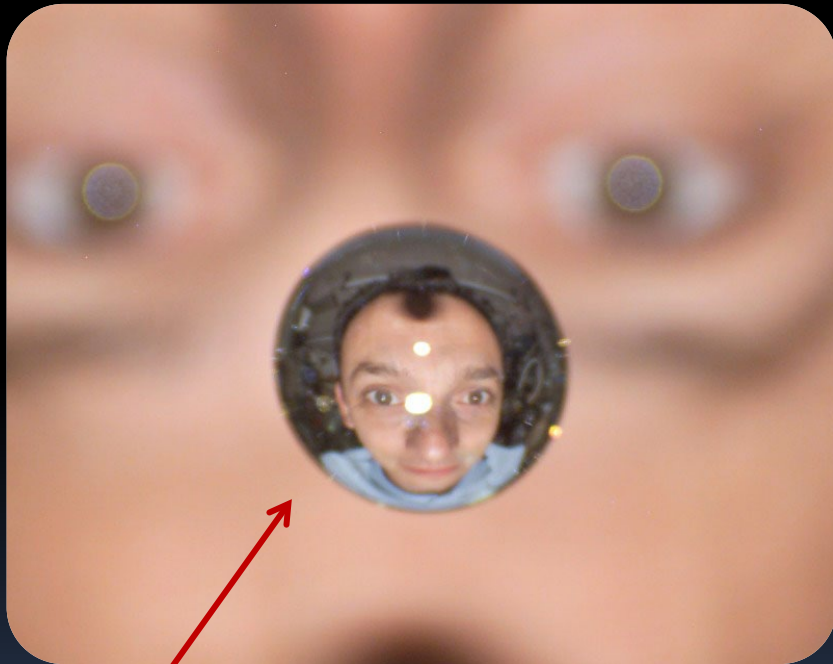
The Higgs Boson

Dilemma: Symmetries work! **VS** Masses are forbidden!

The Higgs Boson

Dilemma: Symmetries work! **VS** Masses are forbidden!

Spontaneous Symmetry Breaking:

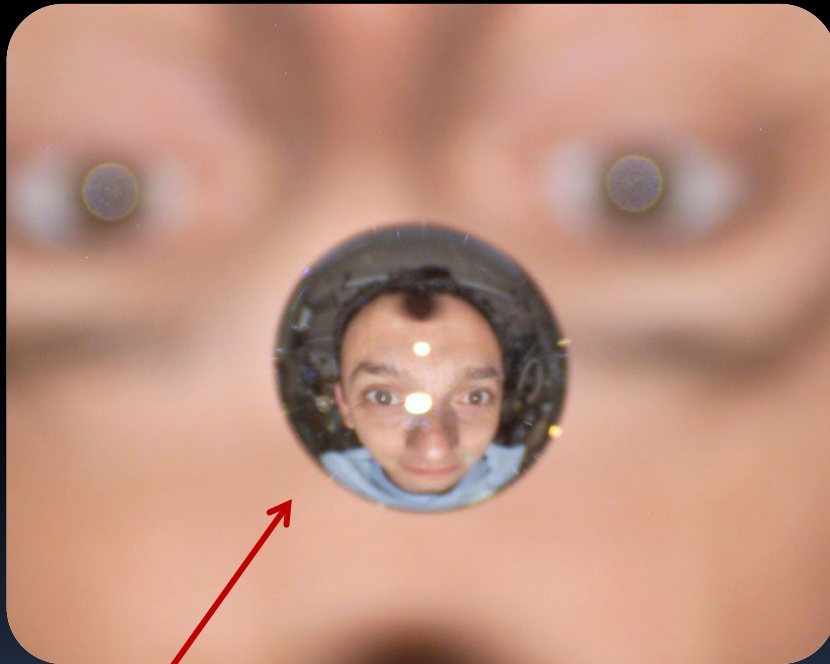


Water droplet (in space)

The Higgs Boson

Dilemma: Symmetries work! **VS** Masses are forbidden!

Spontaneous Symmetry Breaking:



Water droplet (in space)

Cold
→
(lowest energy state)

“Smaller” rotational symmetry

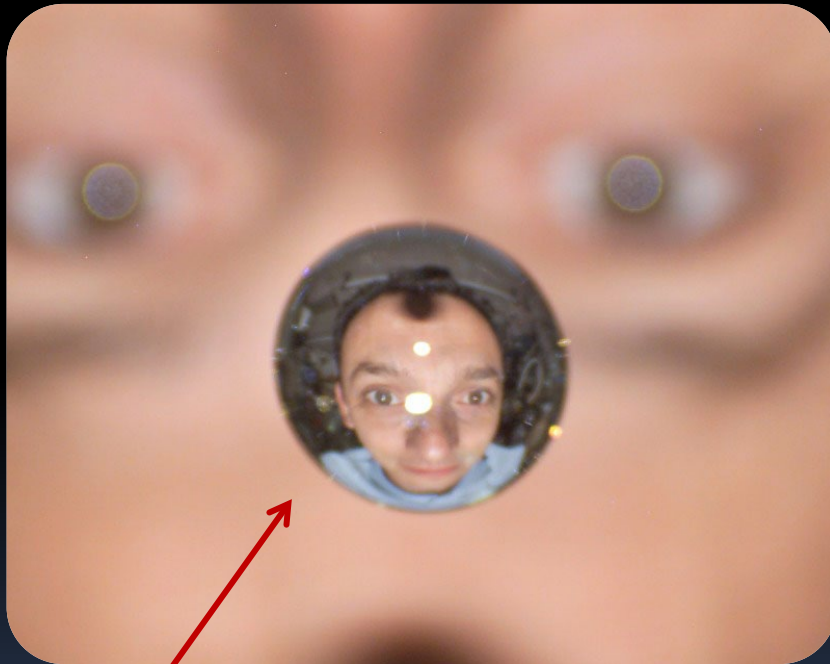


But have we really **LOST** symmetry?

The Higgs Boson

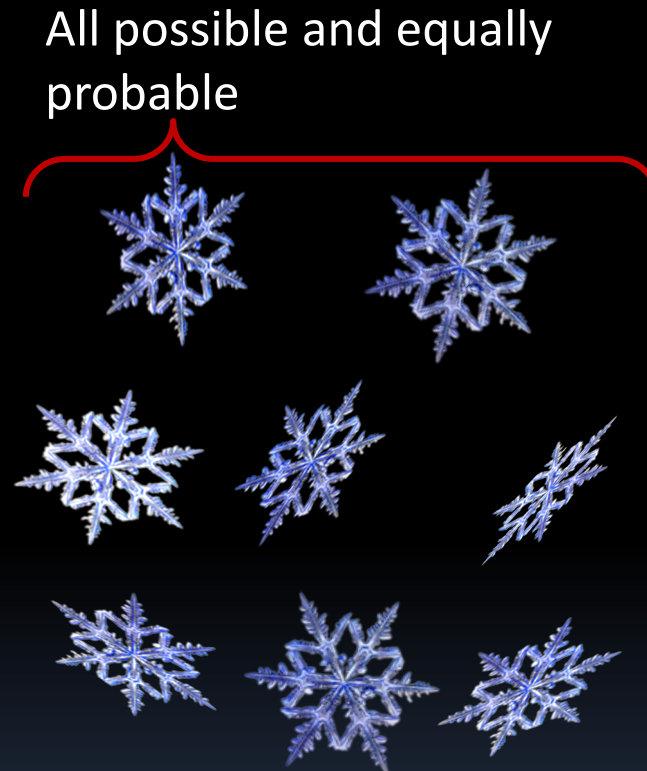
Dilemma: Symmetries work! **VS** Masses are forbidden!

Spontaneous Symmetry Breaking:



Water droplet (in space)

Cold
→
(lowest energy state)



But have we really **LOST** symmetry?

NO! It is just manifested in a more subtle way

The Higgs Boson



Scalar Field:



$SU(2)_L \times U(1)_Y$

Lowest energy state (Vacuum), does not respect symmetry

The Higgs Boson



Scalar Field:

H

$SU(2)_L \times U(1)_Y$

Lowest energy state (Vacuum), does not respect symmetry

The Higgs Boson

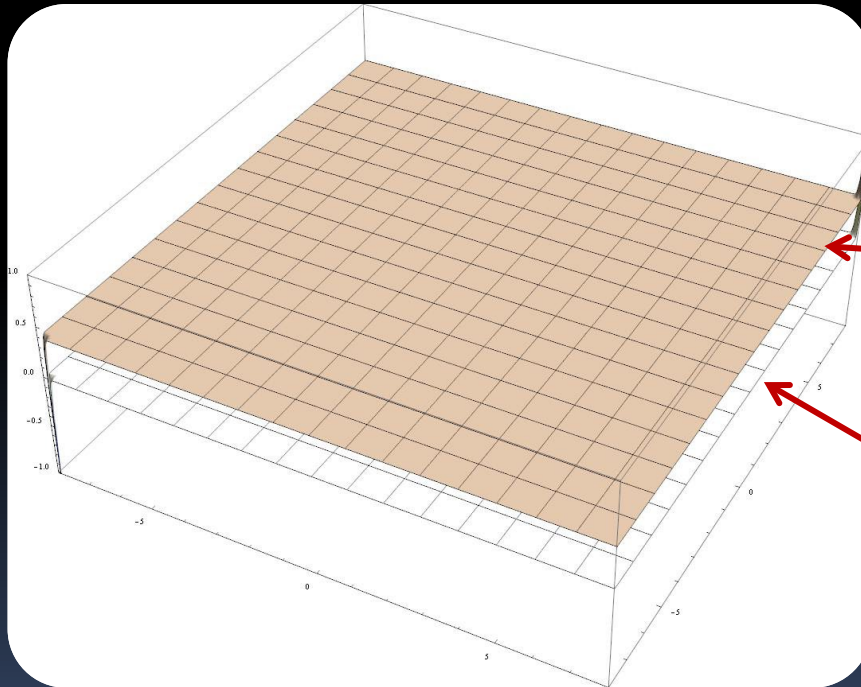
Scalar Field:



$SU(2)_L \times U(1)_Y$

Lowest energy state (Vacuum), does not respect symmetry

Effect: Vacuum Expectation Value (VEV)!



No particles here!

Vacuum of the H field
(symmetry breaking)
(246 GeV)

Normal Vacuum

The Higgs Boson

Any field that interacts with H will constantly “feel” this VEV

Non interacting

Low Interaction



Pictorial “classical” view: even if we allow that “bump” to move at light speed, the effective speed on the interacting case will be smaller than c . Only massive particles can do that.

The Higgs Boson

Any field that interacts with H will constantly “feel” this VEV

Low Interaction

High Interaction



The effect grows with larger interaction (and here the pictorial view fails us).

The Higgs Boson

Theory

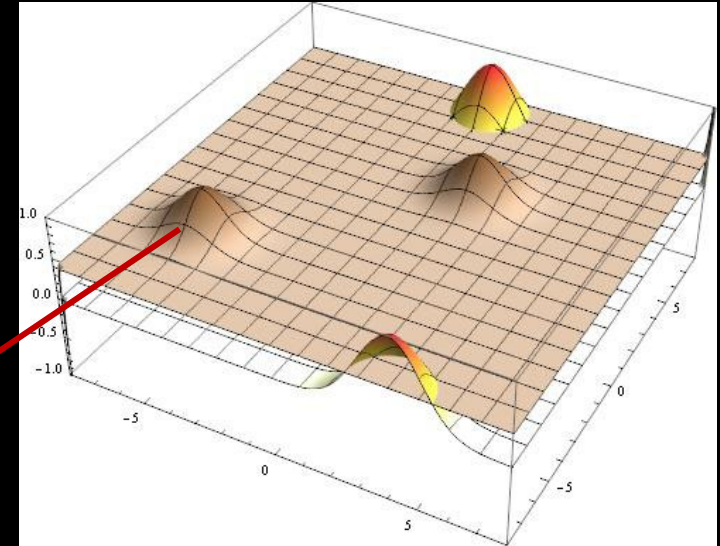
1968

Fermions

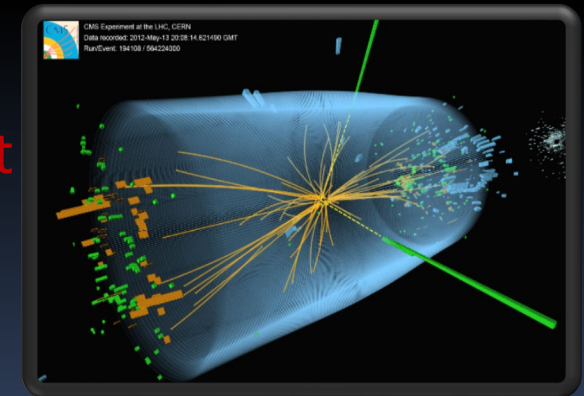
Bosons

$Q = 2/3$	u up	c charm	t top	γ foton
$Q = -1/3$	d down	s strange	b botton	g gluon
$Q = 0$	ν_e electron neutrino	ν_μ muon neutrino	ν_τ tau neutrino	Z Z boson
$Q = -1$	e electron	μ muon	τ tau	W W boson
	1 st gen.	2 nd gen.	3 rd gen.	

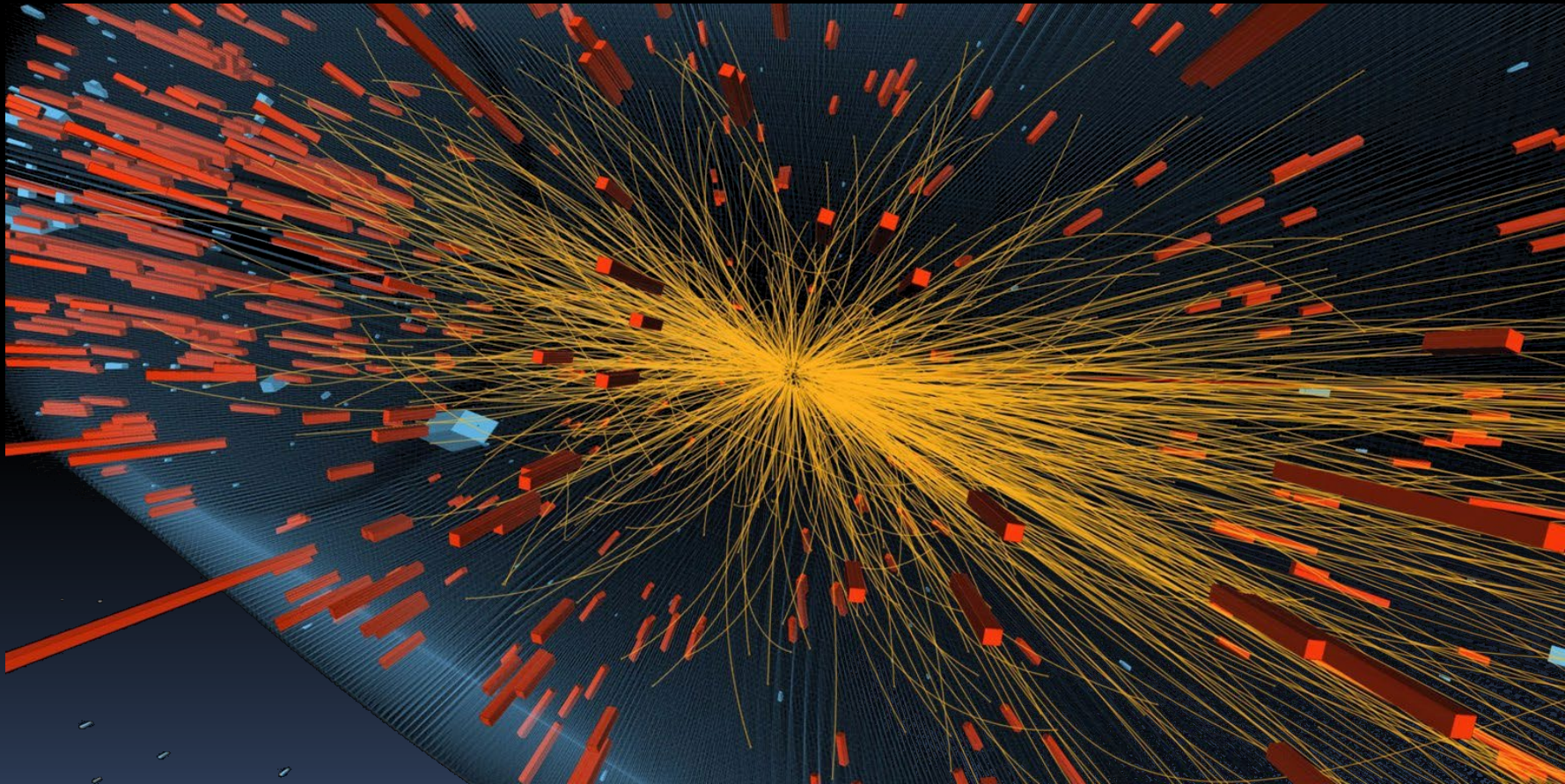
h
Higgs
boson



Experiment
2012

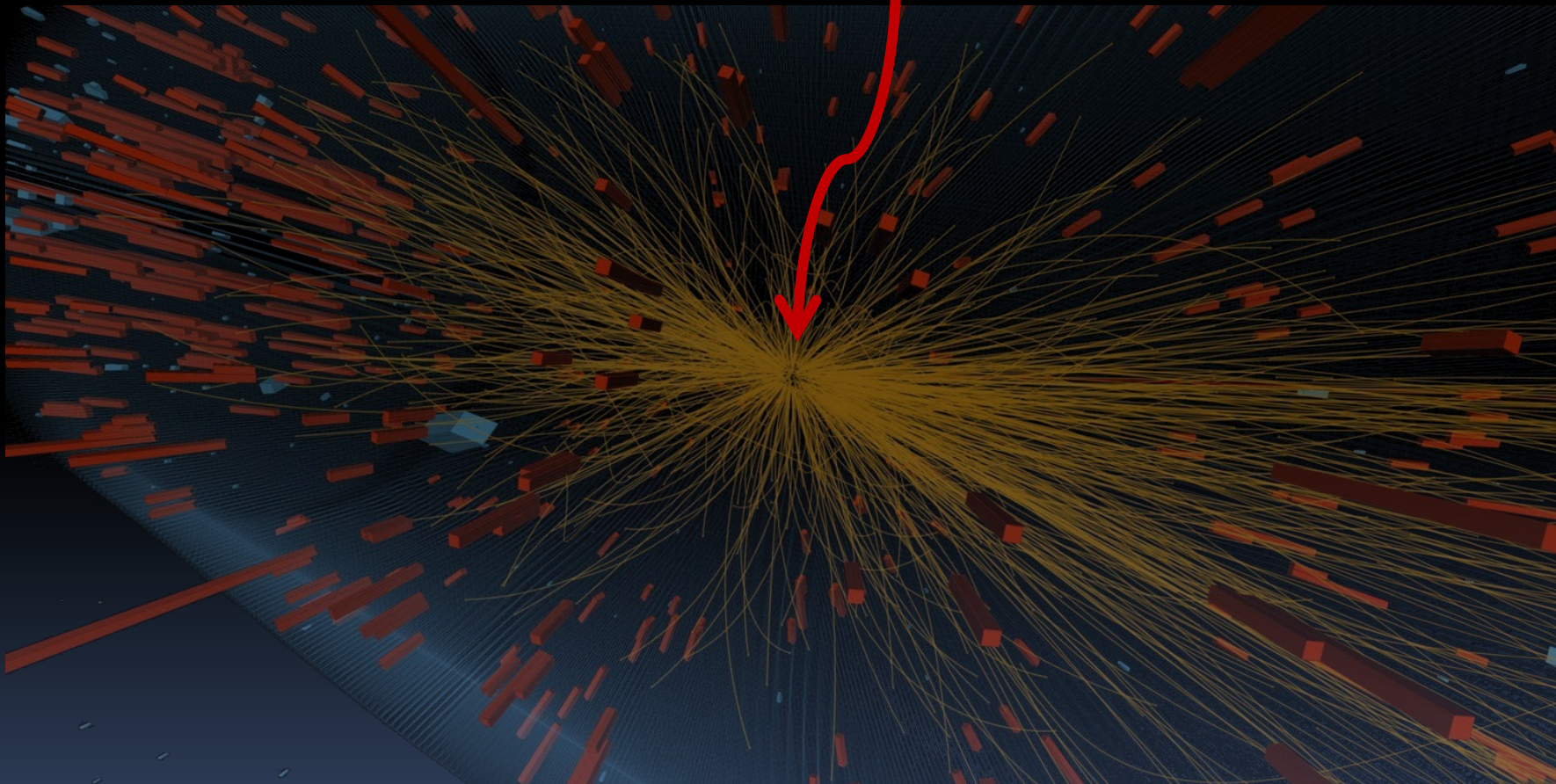


Looking for new particles

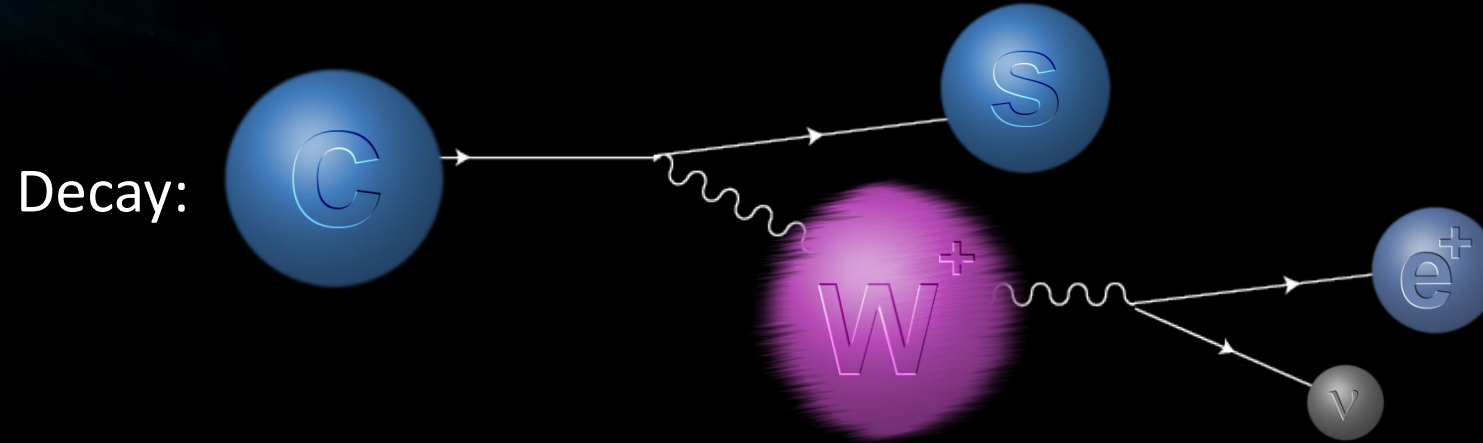


Looking for new particles

How do I know what happened in there?



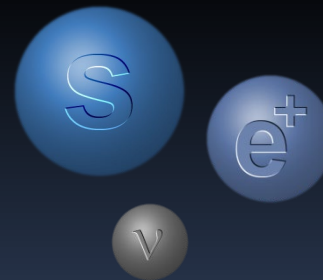
Heavy Particles Decay



Before

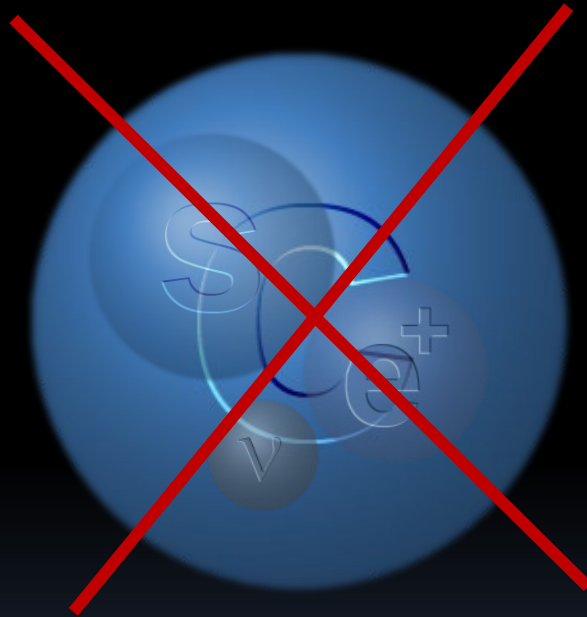


After



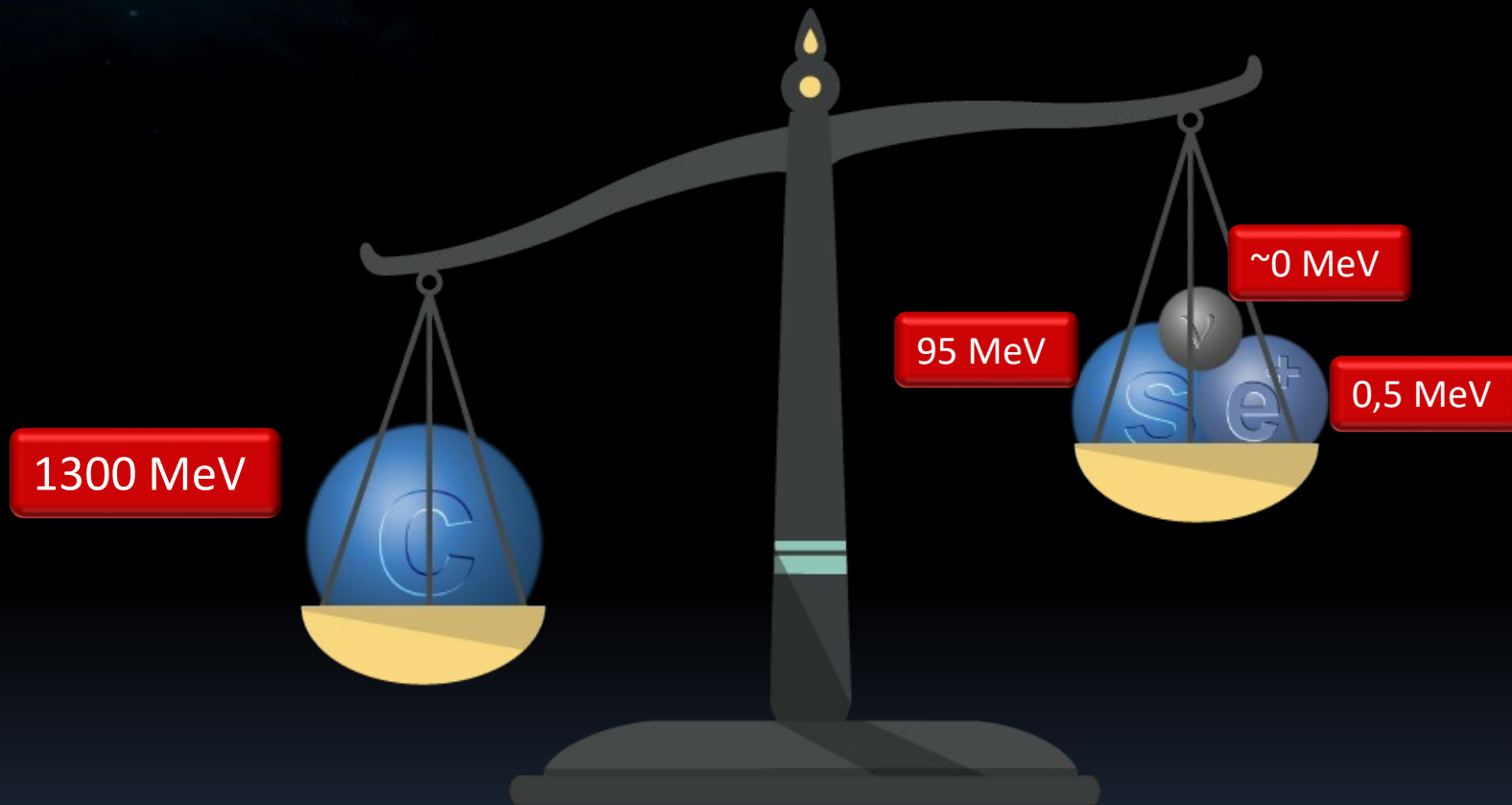
Heavy Particles Decay

Be careful!



Decay does not imply “composition”!

Heavy Particles Decay

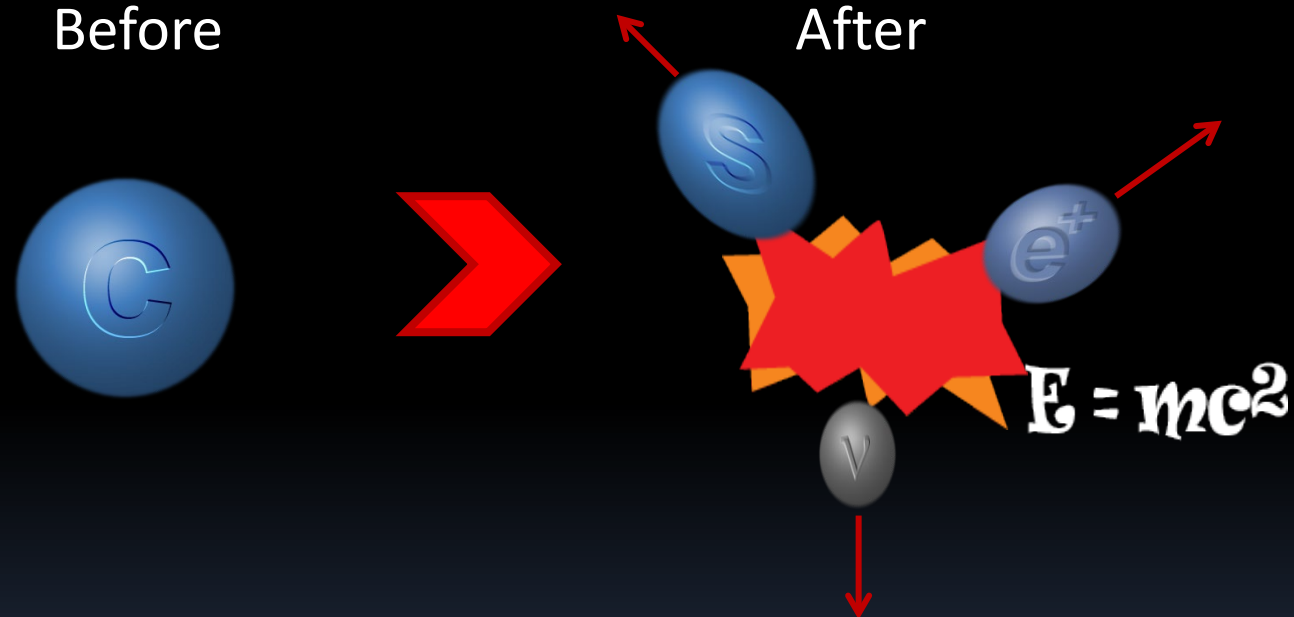


Heavy Particles Decay



This is balanced by kinetic energy for the decay products!

Heavy Particles Decay



If I carefully measure the energy/momentum of the decay products
I can **reconstruct the original state**

Reconstructing the Higgs

As an example, let's look at one of the simplest Higgs **decays**:

125 GeV



Reconstructing the Higgs

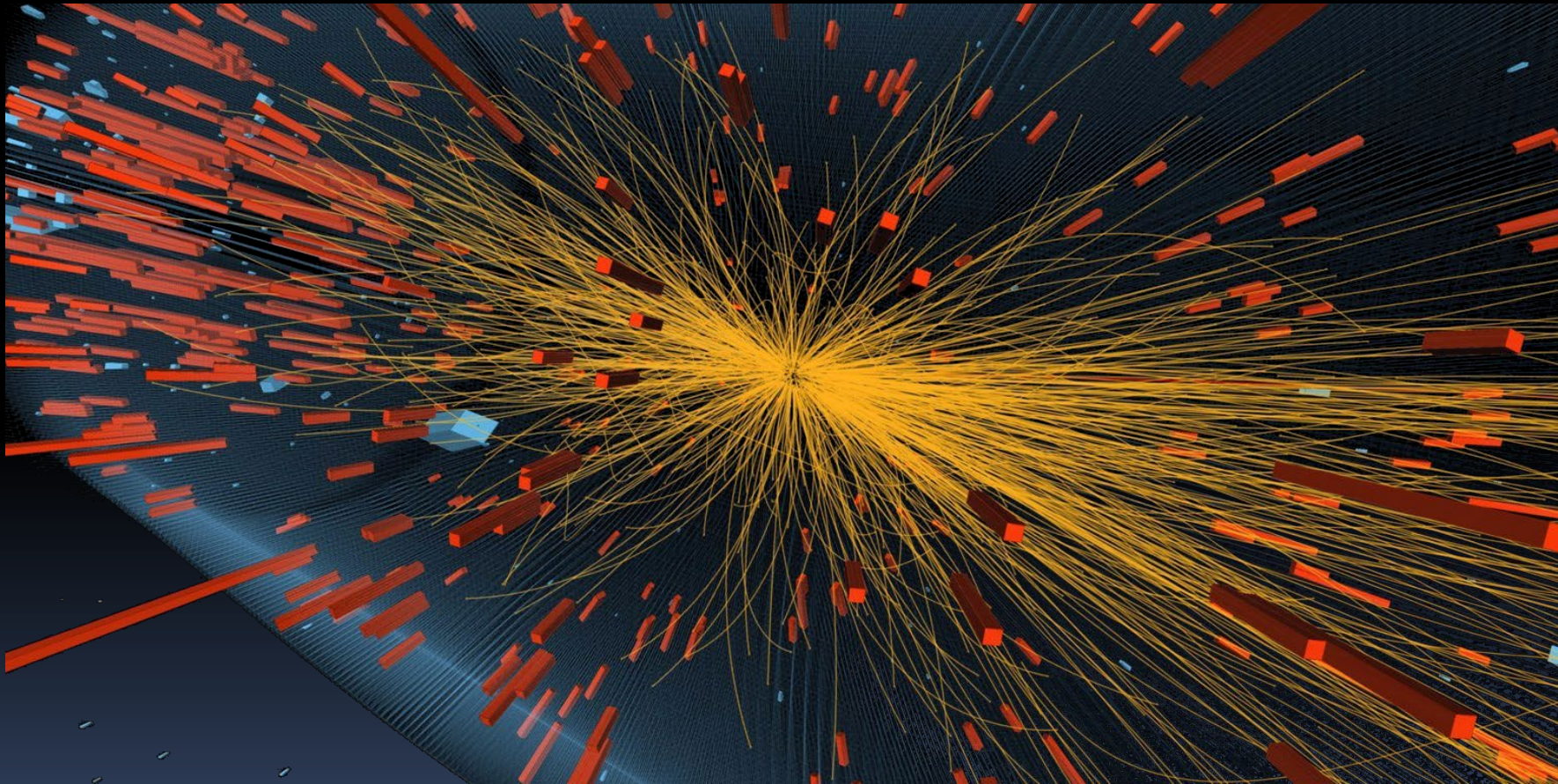
As an example, let's look at one of the simplest Higgs **decays**:



Well, I just have to find two photons with added energy of 125 GeV. **Sounds easy, right?**

Reconstructing the Higgs

What photons would you choose?



Reconstructing the Higgs

What photons would you choose?

Two obvious problems:

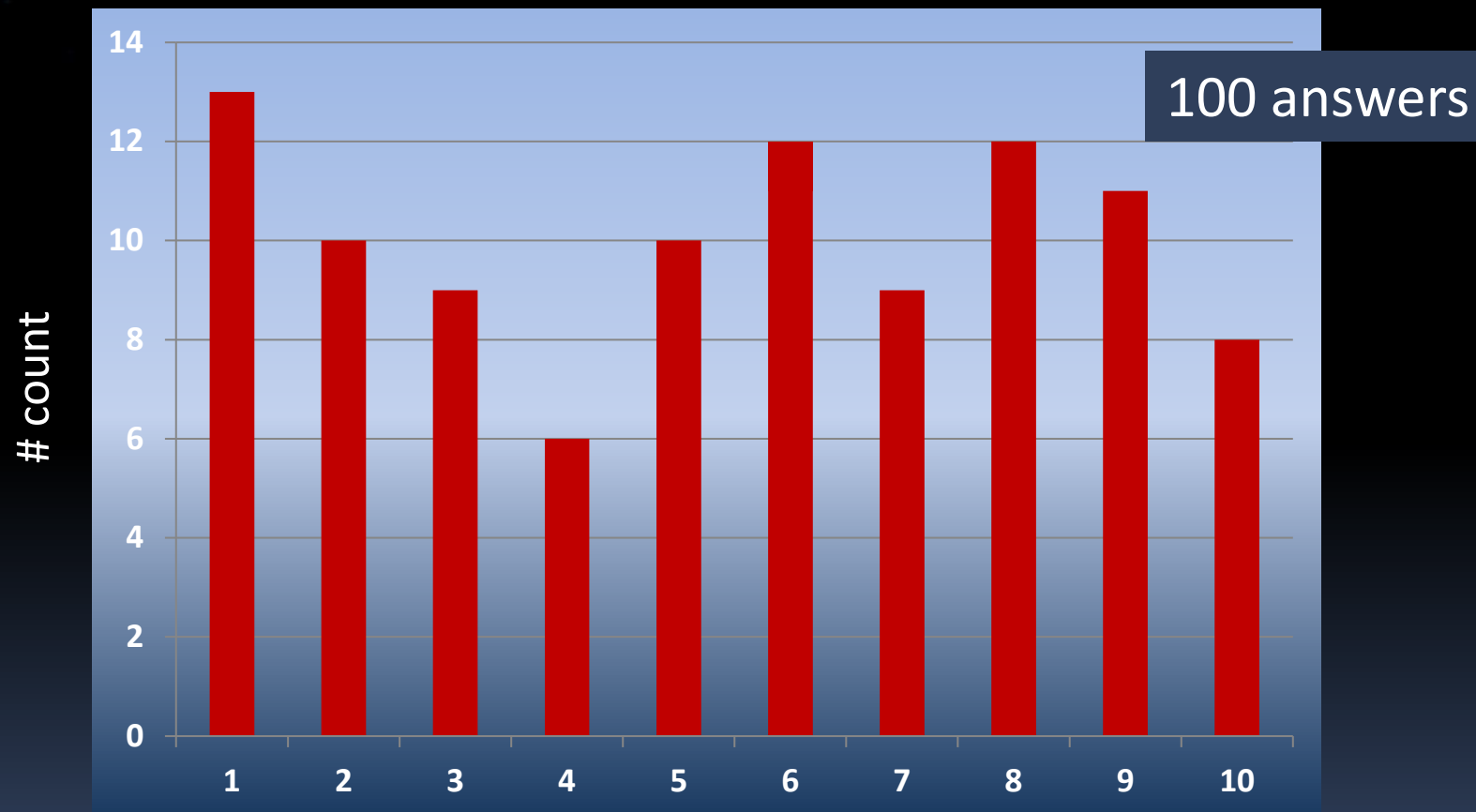
- Photons can add to 125 GeV **by chance**! How do I know is really a Higgs?
- What if I didn't know what the mass is? (in fact **WE DIDN'T** and usually **DON'T**)

Crazy person game

- 100 people in a room, most are able to choose a random number between 1 and 10
- One of them might be crazy, in the sense the he/she always choose the same number
- The crazy person prefers to remain anonymous (and we don't want to anger a crazy person)
- We ask the whole group to choose number in anonymous way, and only do statistics on the resulting sample

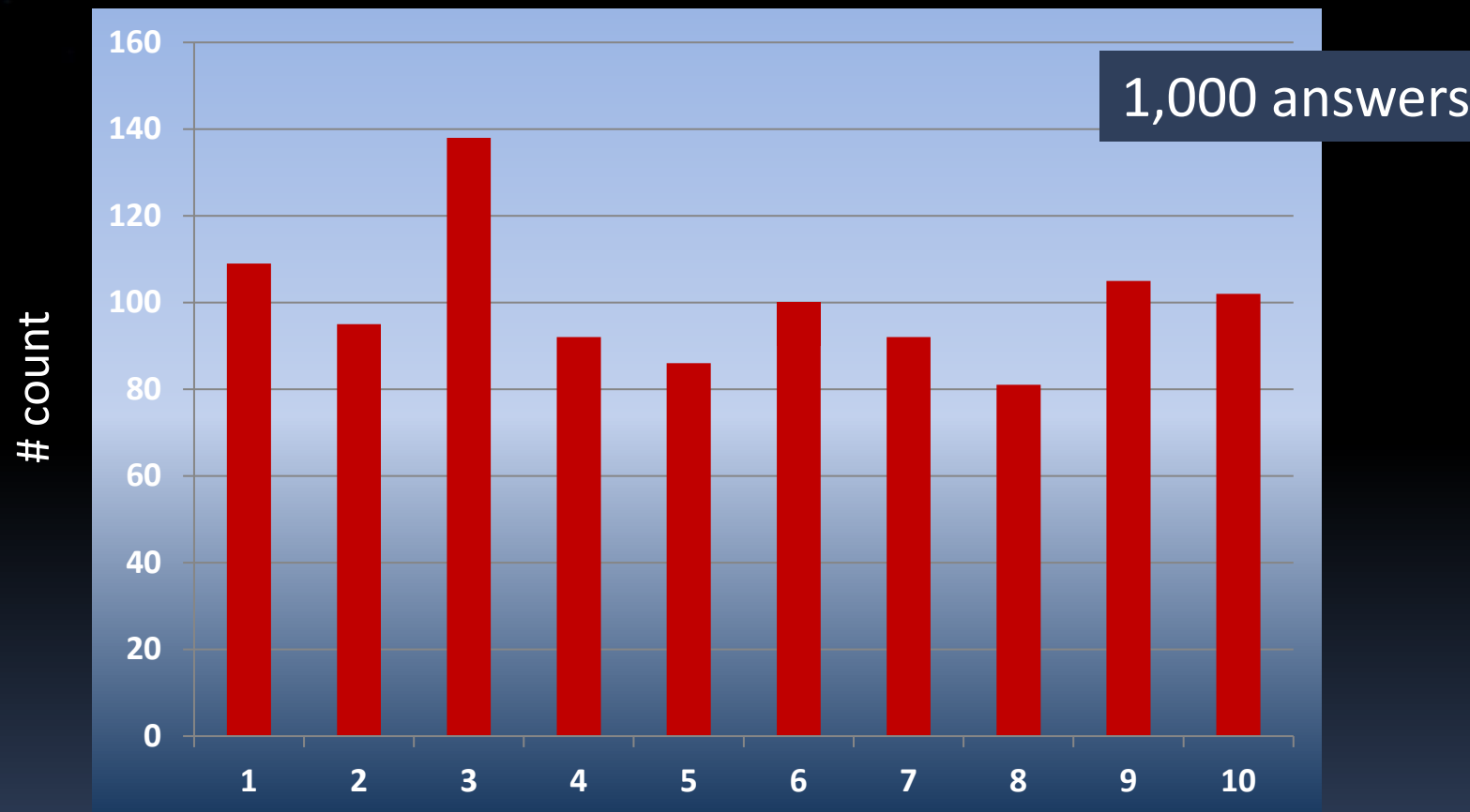
Crazy person game

First sampling (each person deposited one choice):



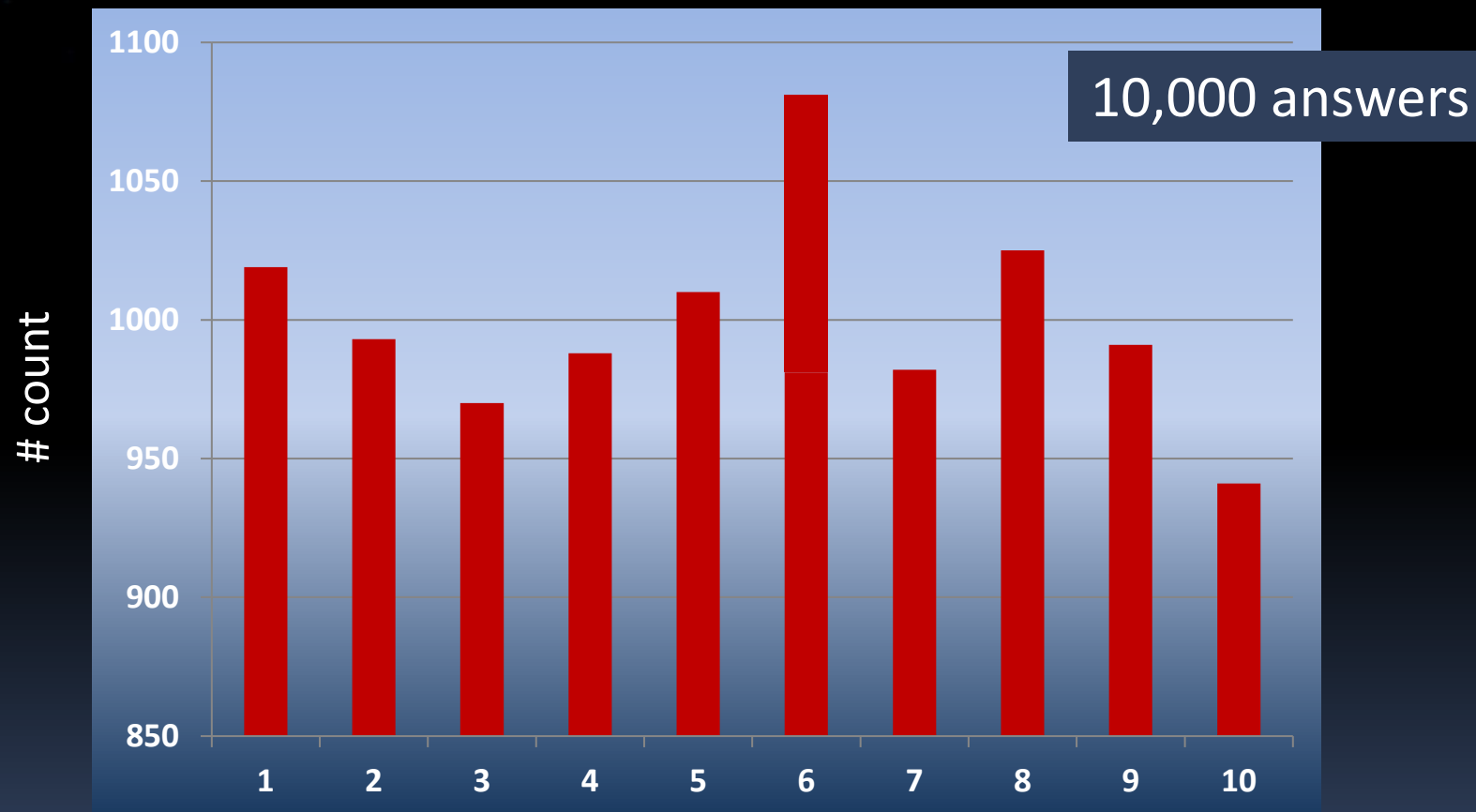
Crazy person game

10 samplings (each person deposited 10 choices):



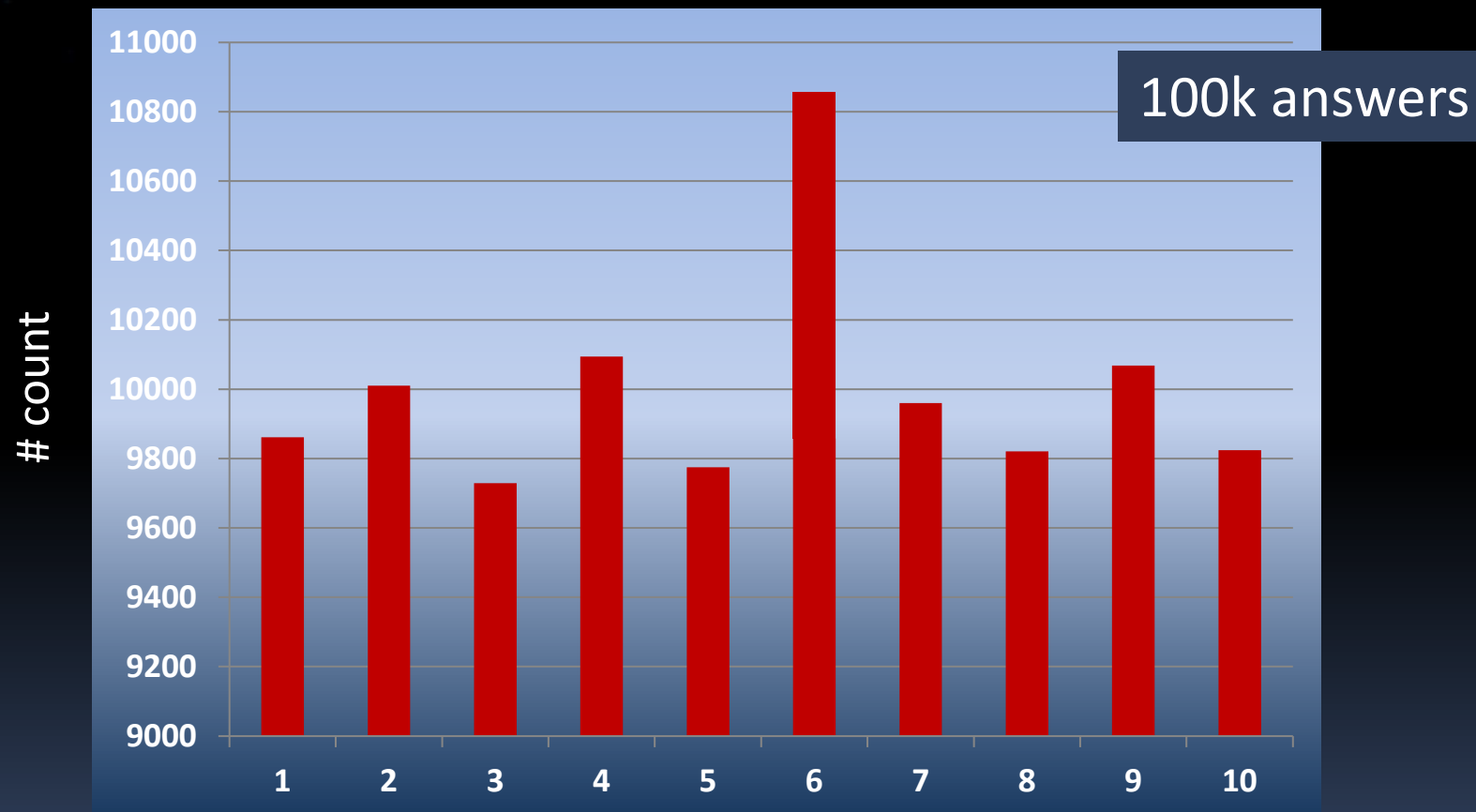
Crazy person game

100 samplings



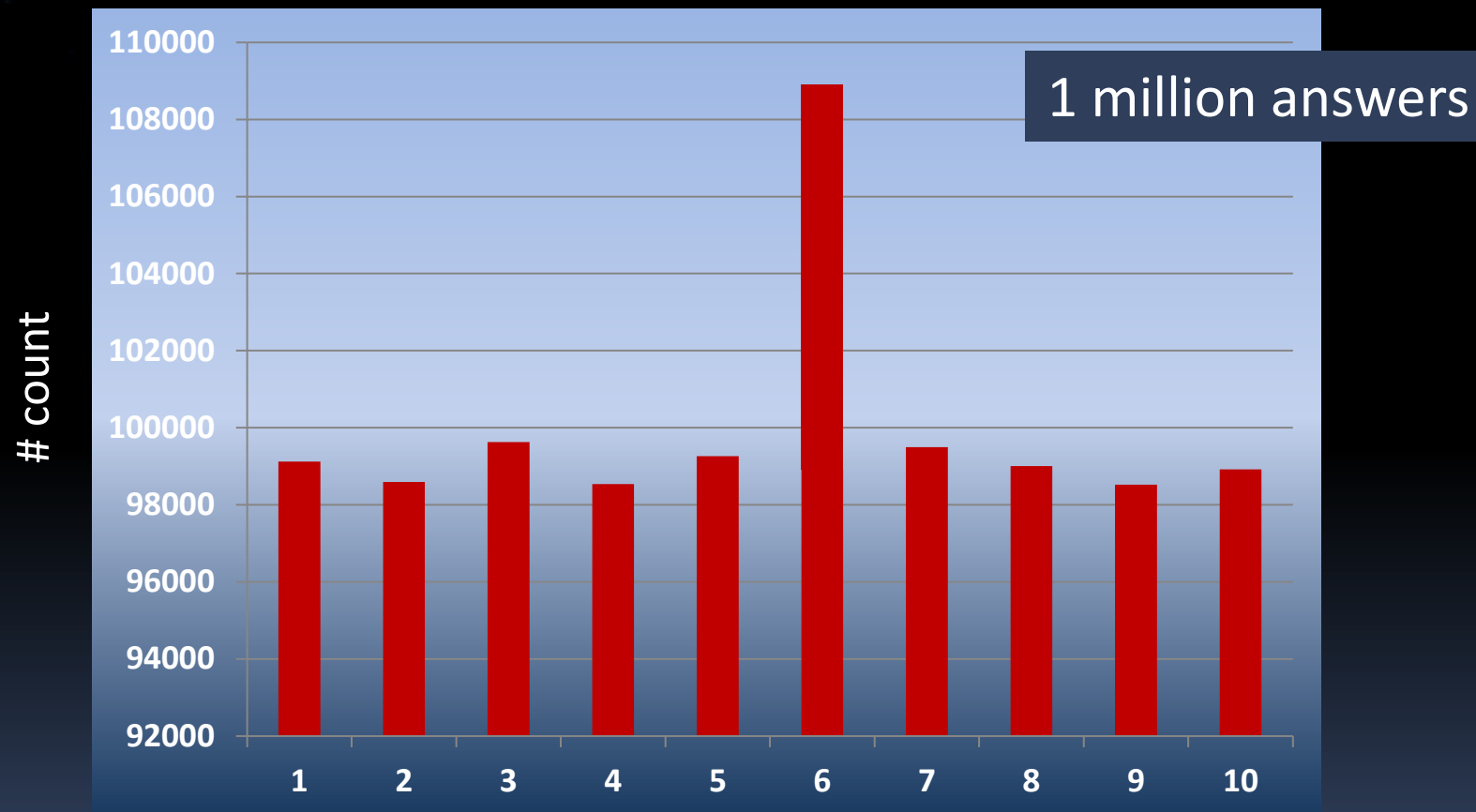
Crazy person game

1,000 samplings



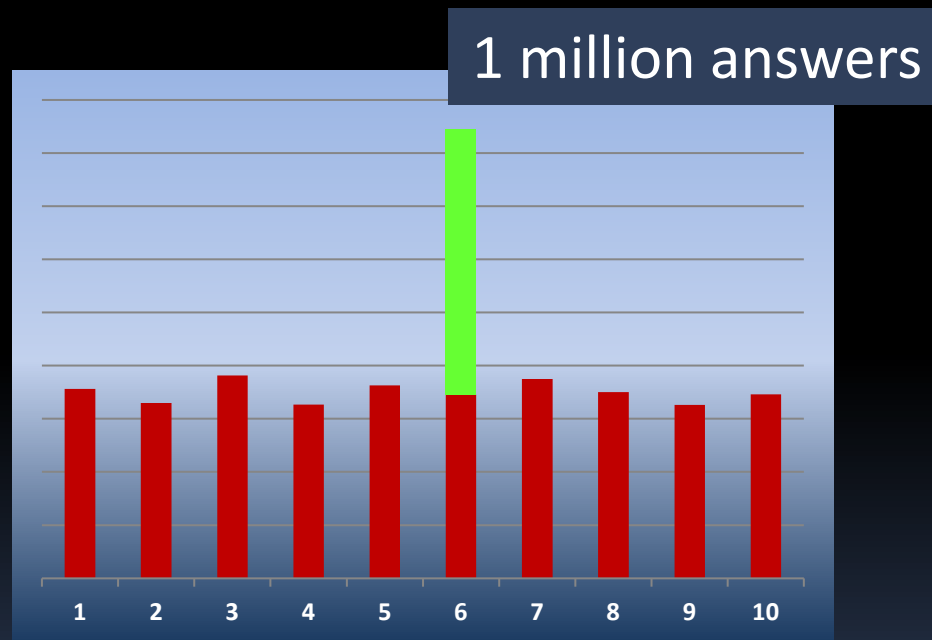
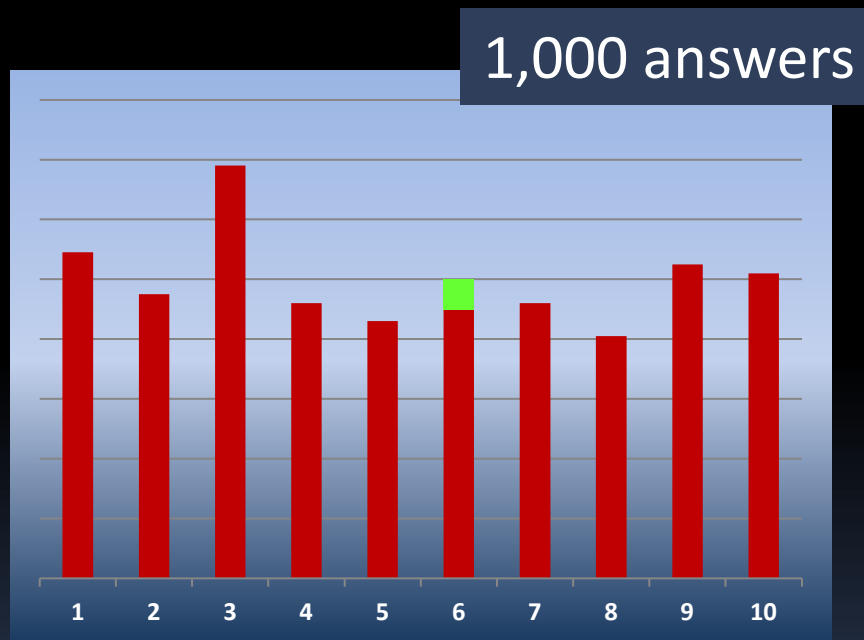
Crazy person game

10,000 samplings

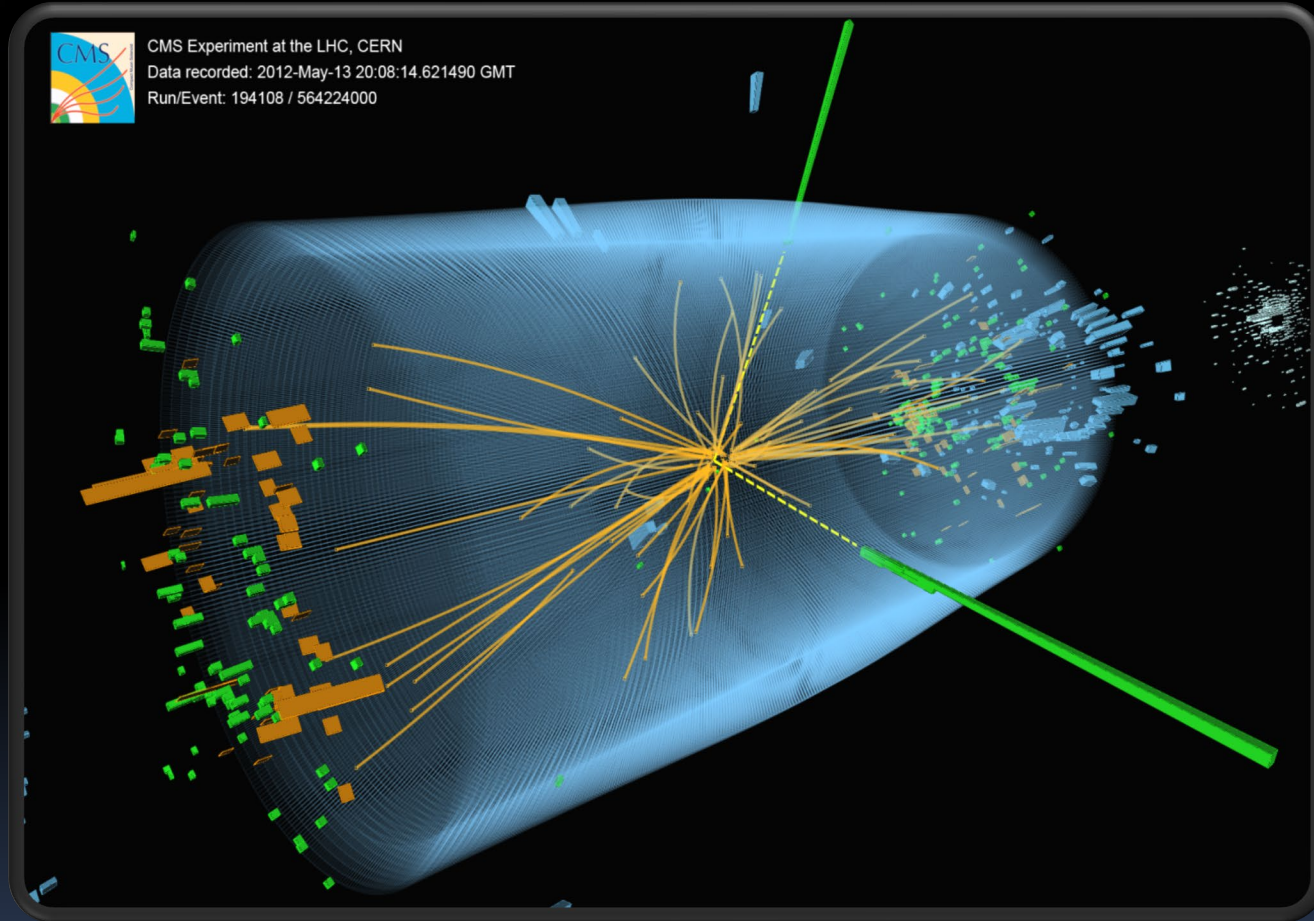
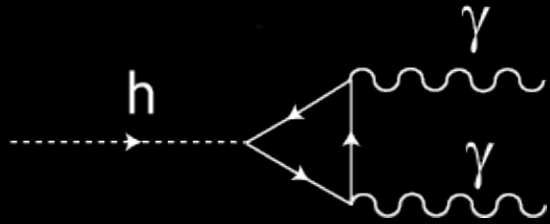


Crazy person game

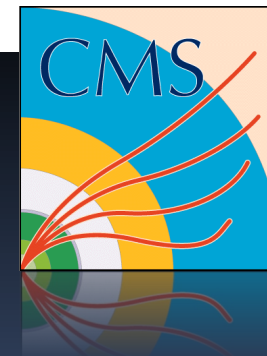
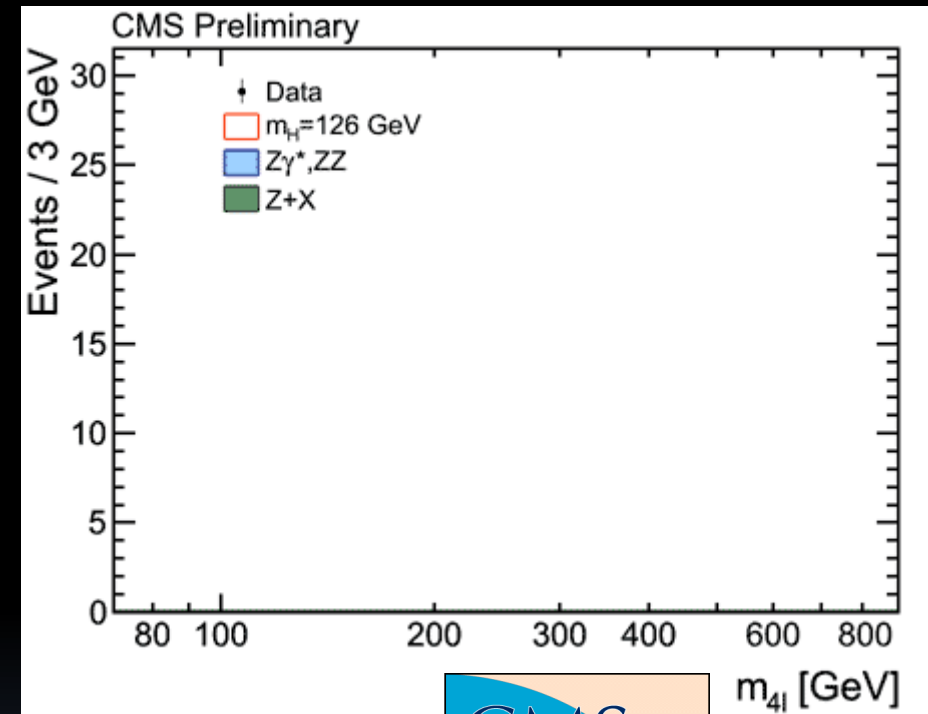
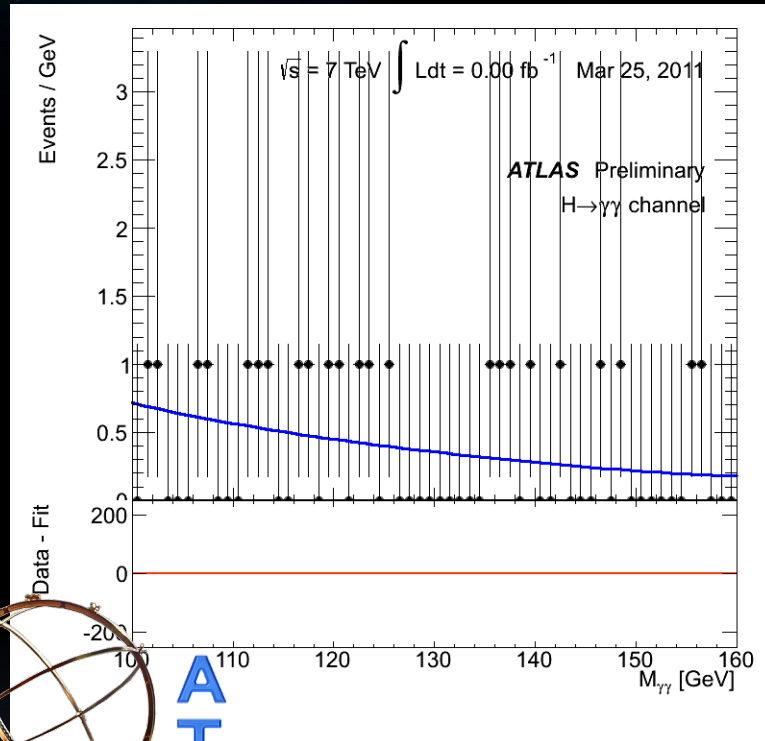
A loucura fica óbvia depois de um número suficientemente grande de repetições



Finding the Higgs Boson

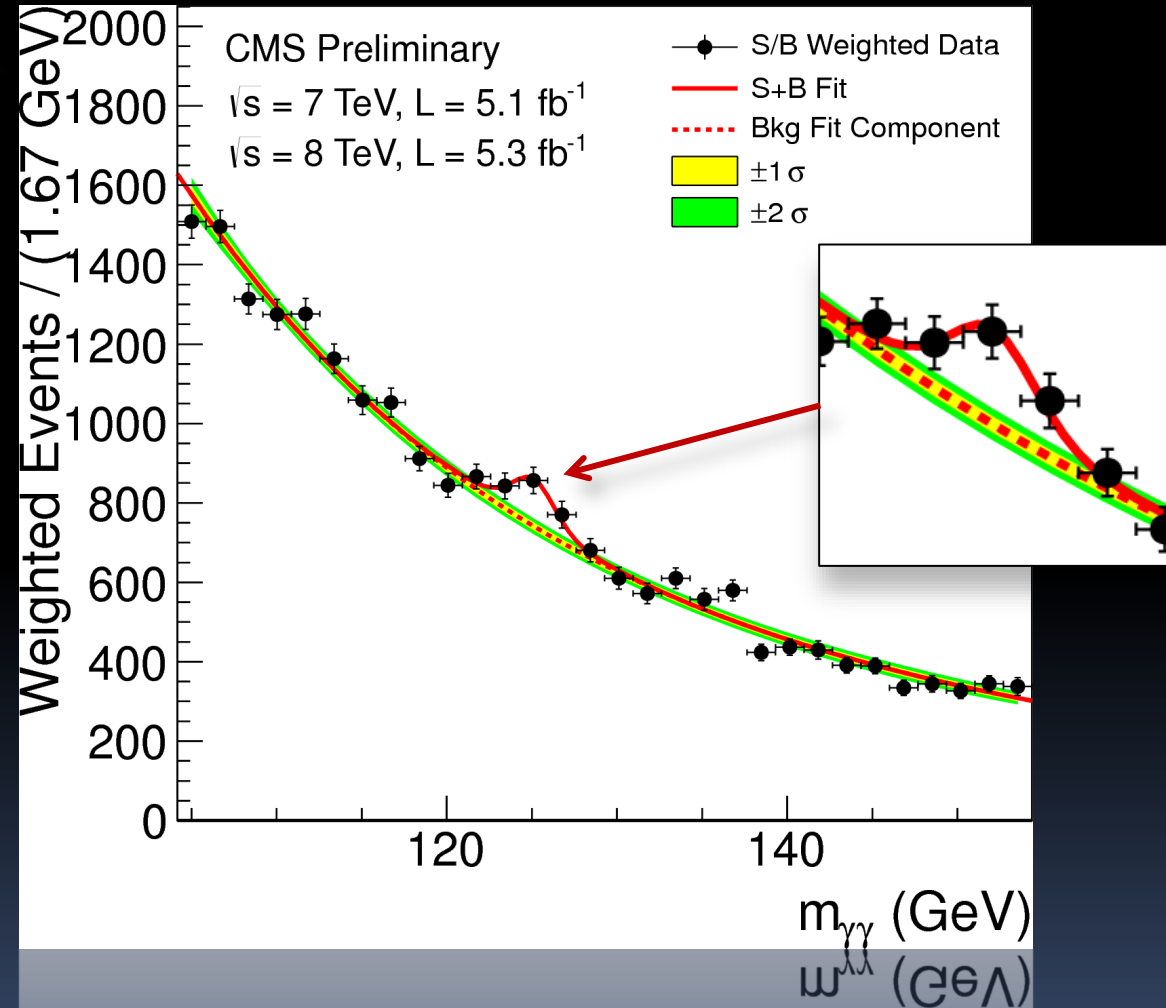
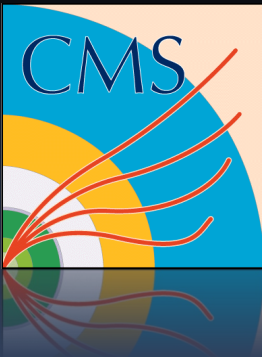
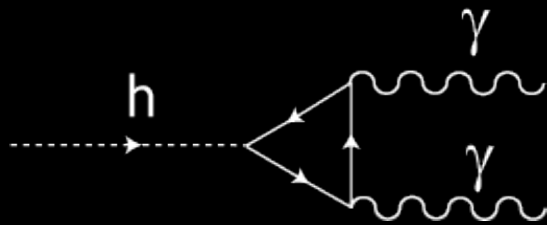


Finding the Higgs Boson



Finding the Higgs Boson

First signs in Dec/2011, and finally in Jul 4th/2012:



Finding the Higgs Boson

First signs in Dec/2011, and finally in Jul 4th/2012:

